

Introdução à Astronomia
Prof. Antônio Kanaan
Aula 4 –
Propriedades da Luz
Instrumentos Astronômicos

Créditos:

A maioria das figuras da aula de hoje vêm de:

1) Chaisson

2) Hypershysics

(<http://hyperphysics.phy-astr.gsu.edu/hbase/hframe.html>)

Informação vem em forma:

1 – luz (ondas eletromagnéticas)

2 – ondas gravitacionais

3 - neutrinos

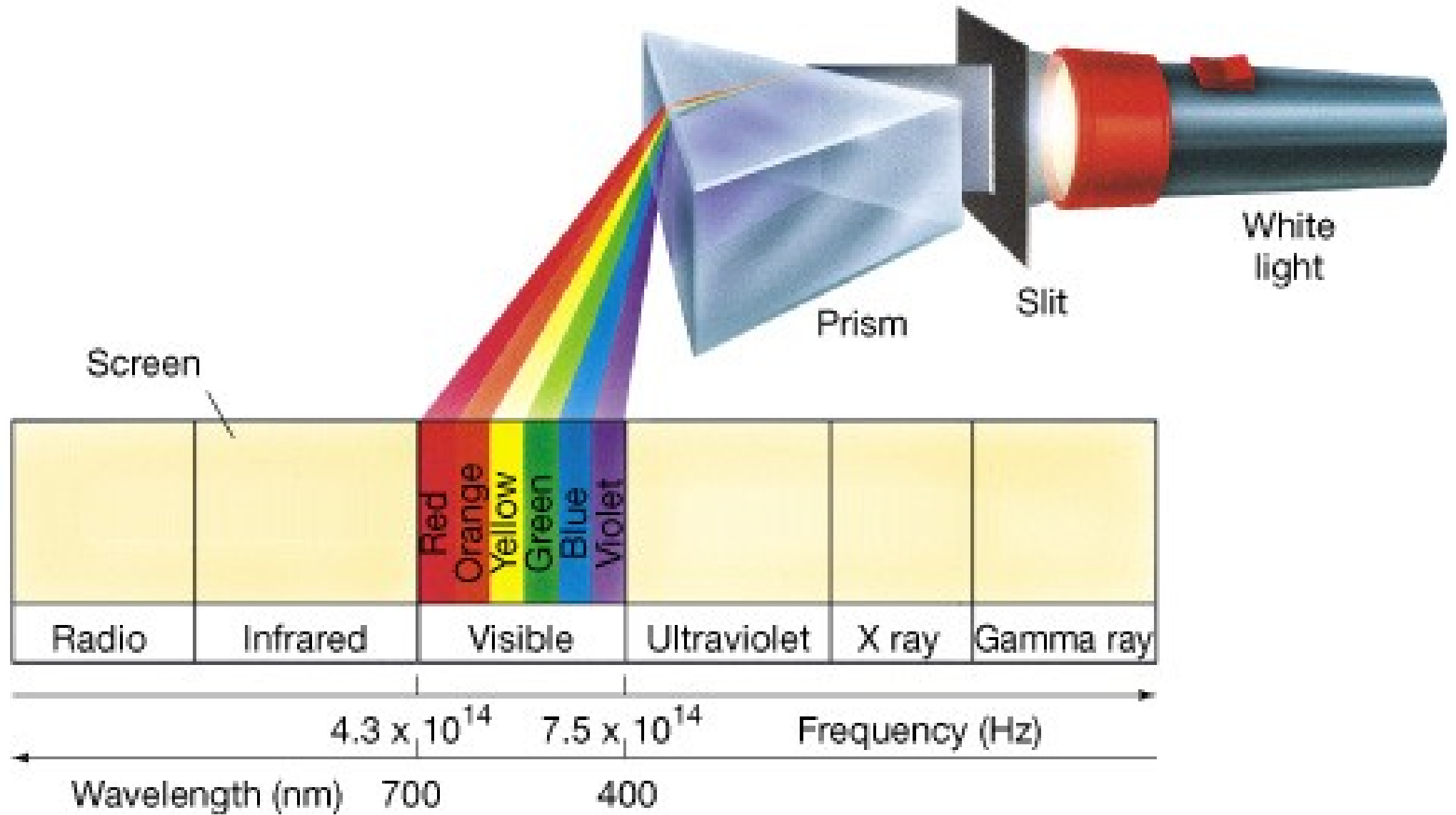
O que é a luz?

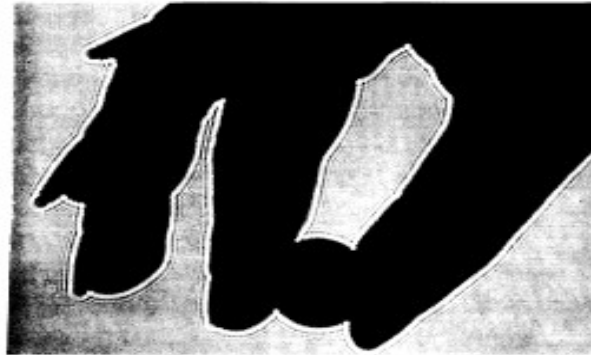
Onda? Partícula? como este é um curso de astronomia e não de física não discutiremos muito sobre a natureza da luz, que é um assunto fascinante.

História curta:

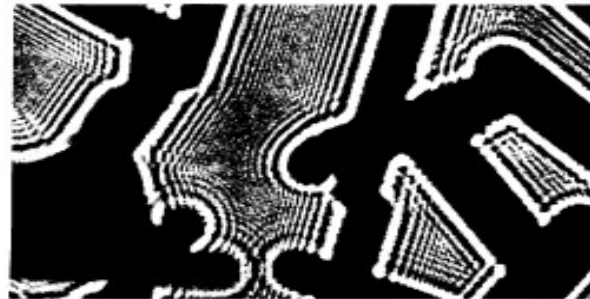
Newton - defensor da teoria corpuscular

Huygens – defensor da teoria ondulatória



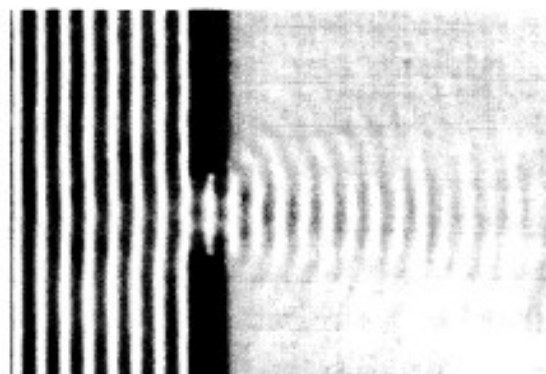


(a)

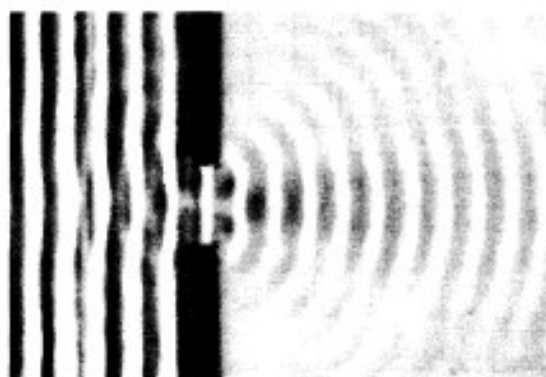


(b)

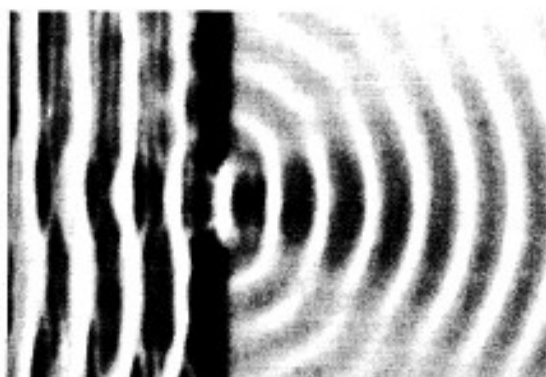
(a) The shadow of Mary's hand holding a dime, cast directly on 4×5 Polaroid A.S.A. 3000 film using a He-Ne beam and no lenses. (Photo by E. H.) (b) Fresnel diffraction of electrons by zinc oxide crystals. (After H. Boersch from *Handbuch der Physik*, edited by S. Flugge, Springer-Verlag, Heidelberg.)



(a)



(b)



(c)

Diffraction through an aperture with varying λ as seen in a ripple tank.
(Photo courtesy PSSC Physics, D. C. Heath, Boston, 1960.)

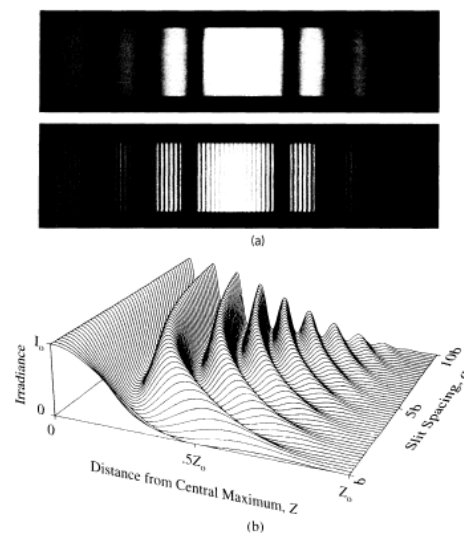
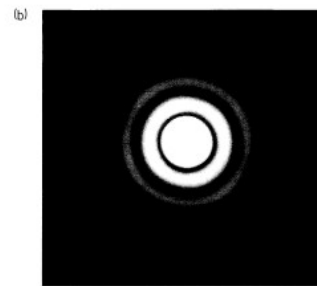
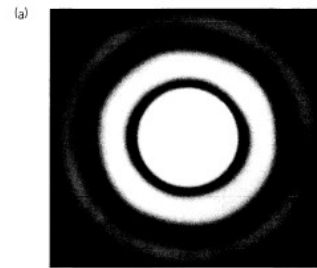
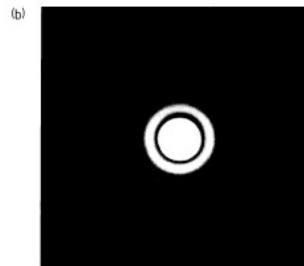
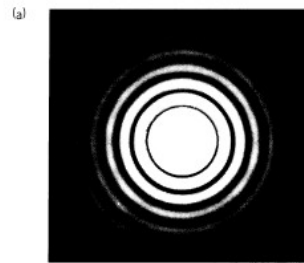


Figure 10.14 Single- and double-slit Fraunhofer patterns. (a) Photographs taken with monochromatic light. The faint cross-hatching arises entirely in the printing process. (Photos courtesy M. Cagnet, M. Francon, and J. C. Thieerr: *Atlas optischer Erscheinungen*, Berlin-Hedelberg-New York: Springer, 1962.) (b) When the slit spacing equals b , the two slits coalesce into one (of width $2b$) and the single-slit pattern appears—that's the first curve closest to you. The farthest curve corresponds to the two slits separated by $a = 10b$. Notice that the two-slit patterns all have their first diffraction minimum at a distance from the central maximum of Z_0 . Note how the curves gradually match Fig. 10.13b as the slit width b gets smaller in comparison to the separation a . (Reprinted from "Graphical Representations of Fraunhofer Interference and Diffraction" *Am. J. Phys.*, 62, 6, (1994), with permission of A. B. Bartlett, University of Colorado and B. Mechtly, Northeast Missouri State University and the American Association of Physics Teachers.)



Airy rings using (a) a 0.5-mm hole diameter and (b) a 1.0-mm hole diameter. (Photo by E. H.)



(a) Airy rings—long exposure (1.5-mm hole diameter). (b) Central Airy disk—short exposure with the same aperture. (Photos by E. H.)

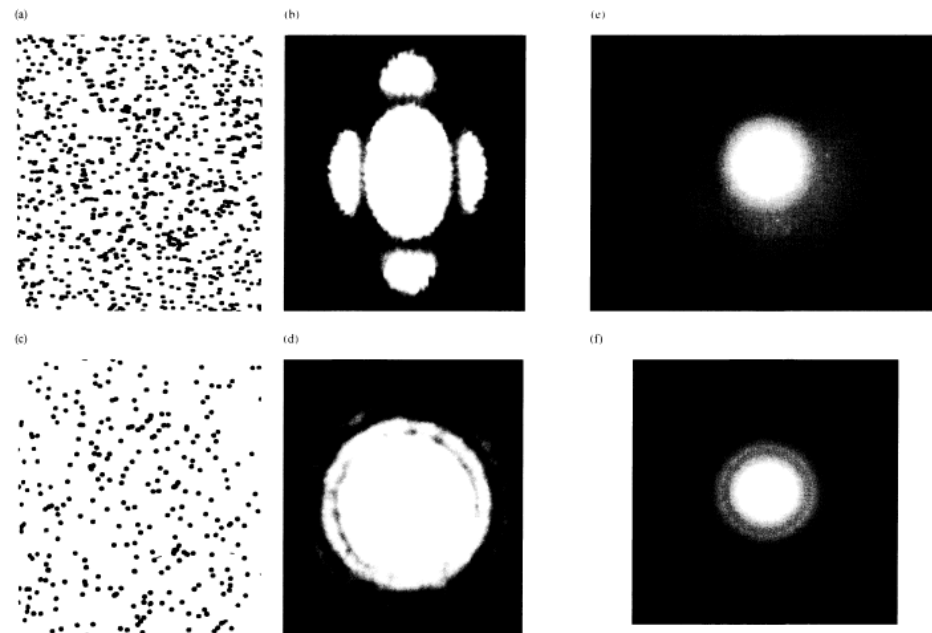
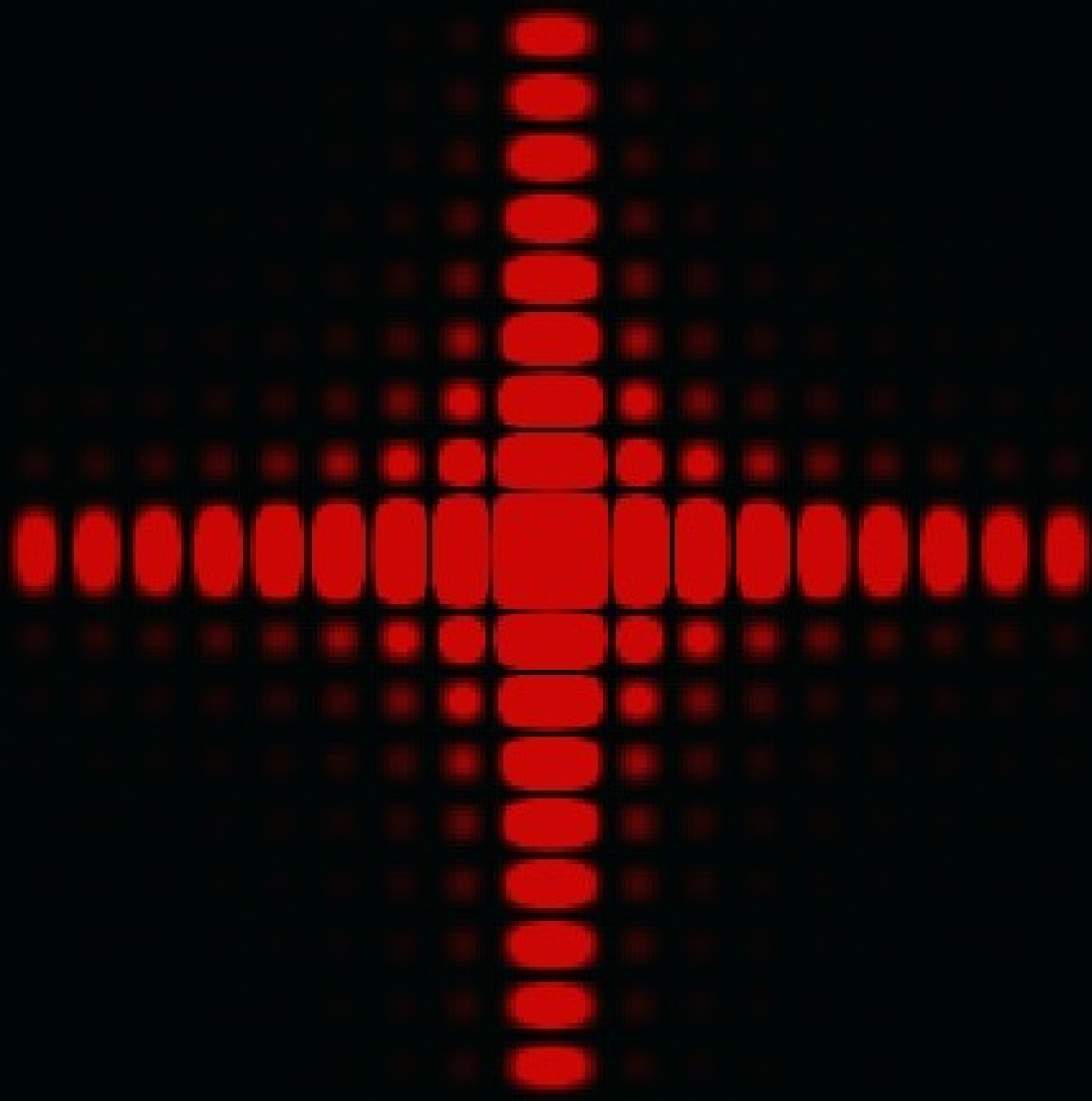


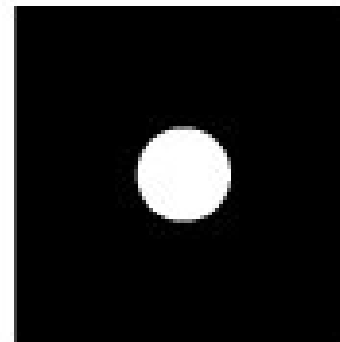
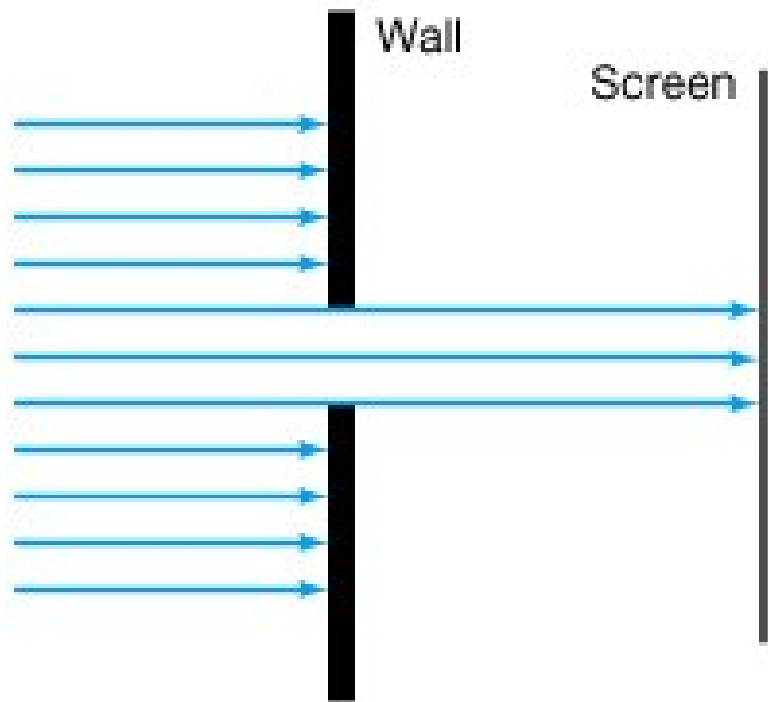
Figure 10.33 (a) A random array of rectangular apertures. (b) The resulting white-light Fraunhofer pattern. (c) A random array of circular apertures. (d) The resulting white-light Fraunhofer pattern. (Photos courtesy The Ealing Corporation and Richard B. Hoover.) (e) A candle flame viewed through a fogged piece of glass. The spectral colors are visible as concentric rings. (Photo by E. H.) (f) A similar colored ring system created by viewing a white-light point source through a glass plate covered with transparent spherical lycopodium spores. (Photo by E. H.)



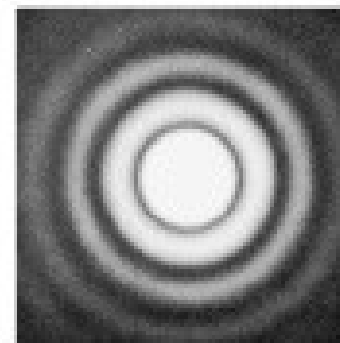
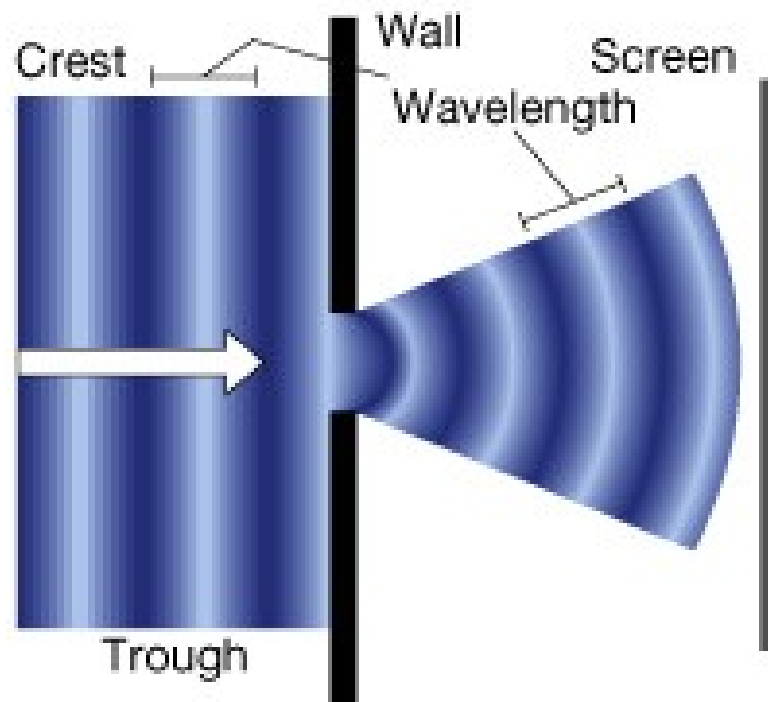
Difração de uma onda de água em um tanque

Difração da luz de
um laser ao passar
por um orifício
quadrado.





Sharp-edged shadow



Fuzzy shadow

O éter:

se onda se propaga como?

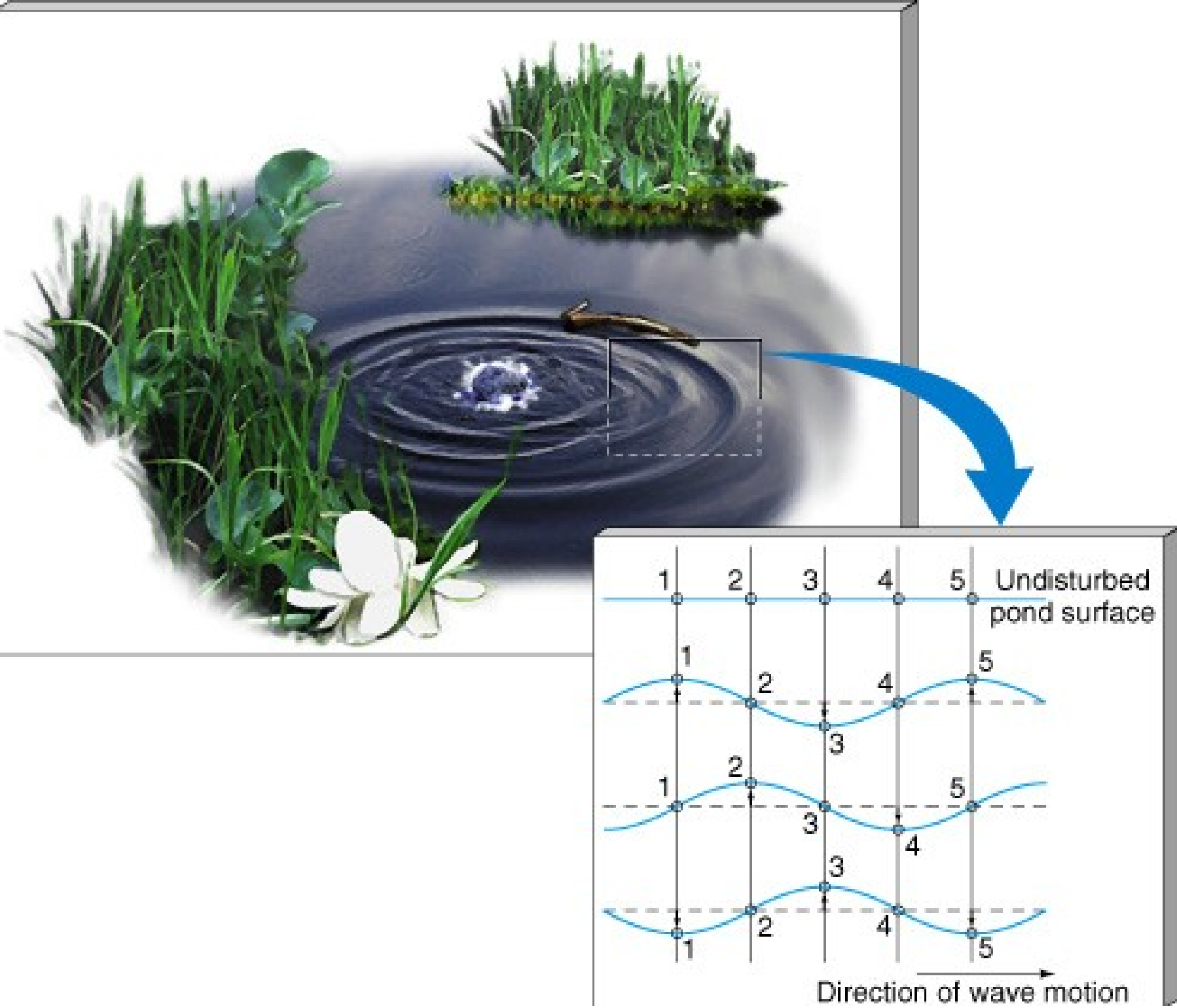
inventou-se o éter.

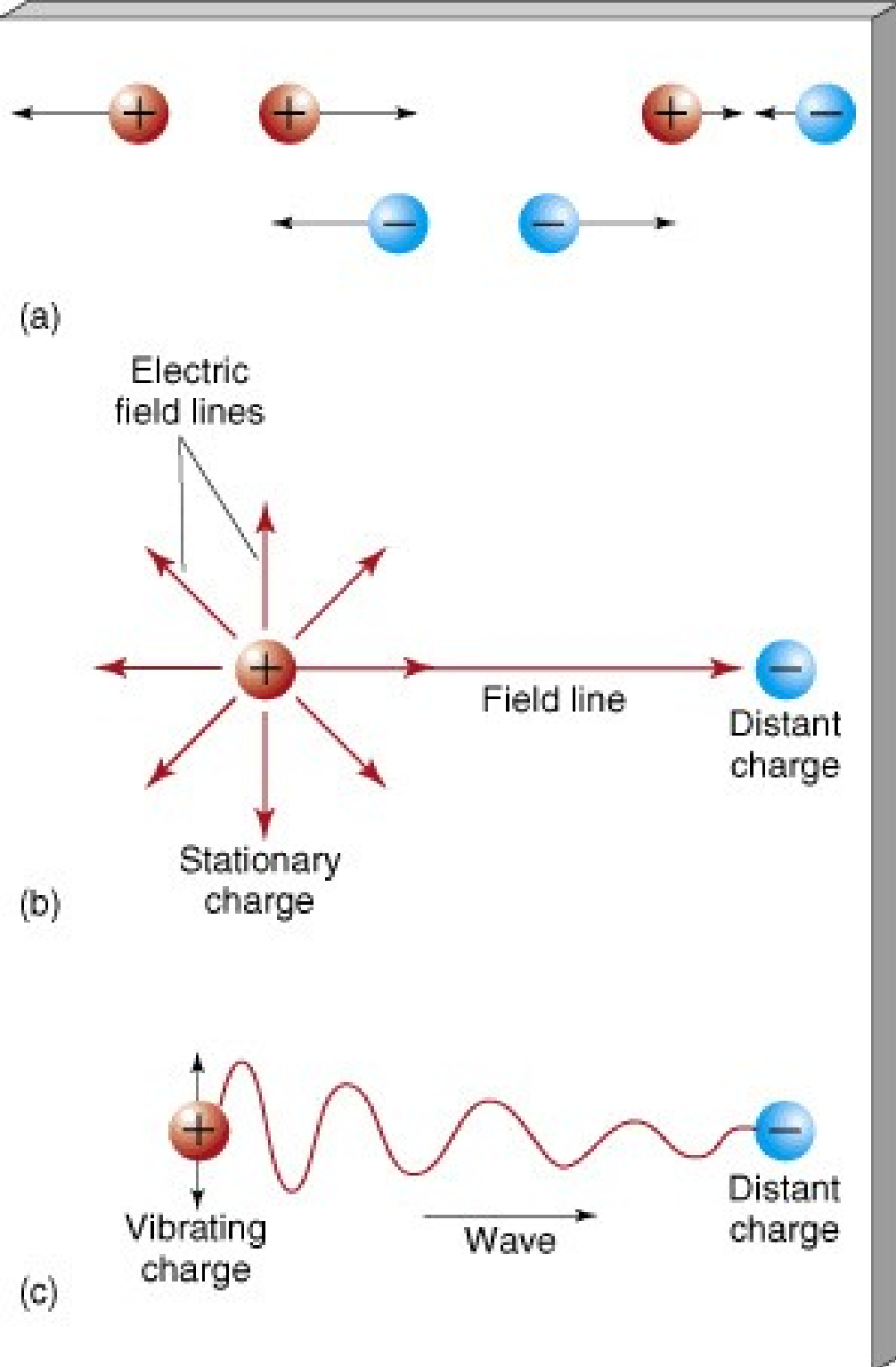
por fim se provou que o éter não existe (experiência de Michelson-Morley)

“provou-se” mais ainda: que c =constante

e por fim acabou se provando o que “não se queria”: que a luz também é uma partícula, daí começa a grande revolução da mecânica quântica.

efeito fotoelétrico:





O éter:

se onda se propaga como?

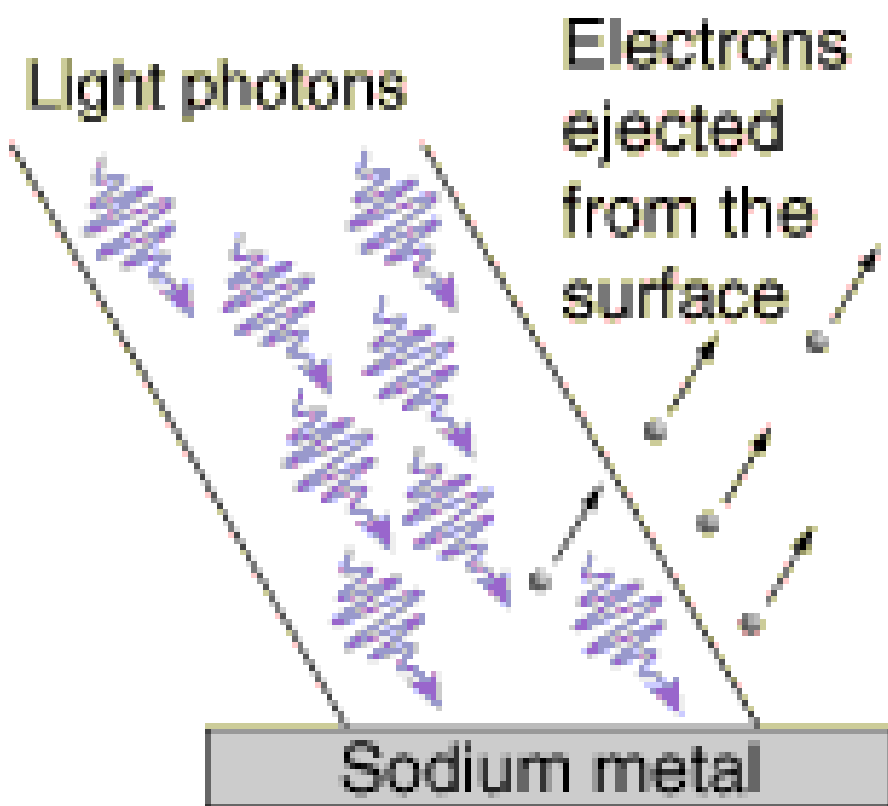
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efeito fotoelétrico:

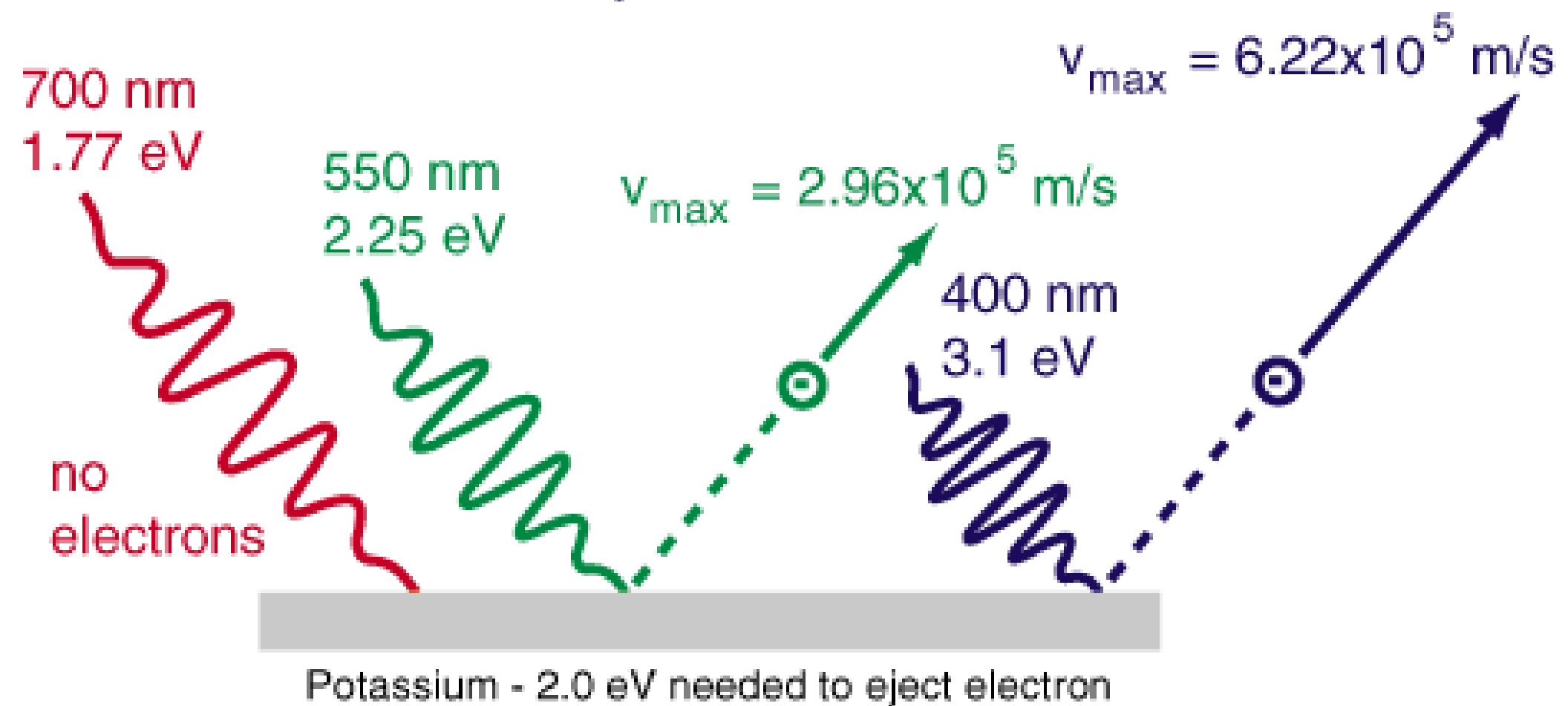


Photon energy

$$E = h\nu$$

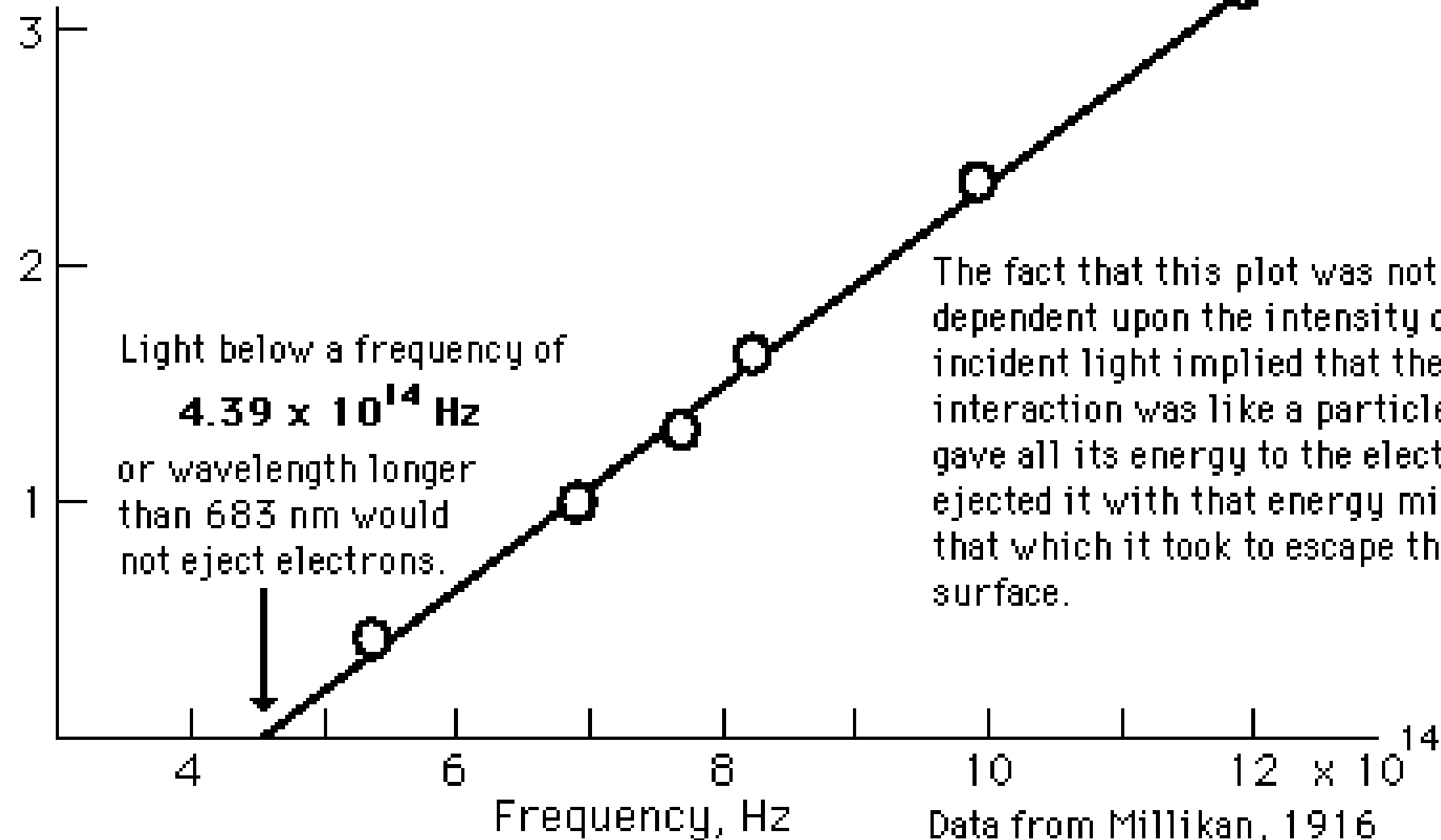
explains the experiment
and shows that light
behaves like particles.

$$E_{\text{photon}} = h\nu$$



Photoelectric effect

Maximum photoelectron kinetic energy in eV



Só há geração de elétrons quando a luz tem frequência maior que um determinado patamar

AM (540-1650 KHz) FM (88-108 MHz) Microwave

Radio

1 GHz 100 GHz

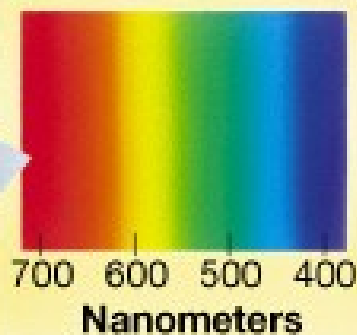
Infrared
far near
100 microns 1

Visible

Ultraviolet
near far

X rays
"Soft" "Hard"

Gamma rays



Frequency
(Hertz)

10^3 10^5 10^7 10^9 10^{11} 10^{13} 10^{15} 10^{17} 10^{19} 10^{21} 10^{23}

Wavelength
(meters)

10^4 10^2 1 10^{-2} 10^{-4} 10^{-6} 10^{-8} 10^{-10} 10^{-12} 10^{-14}

Scale

Mount Everest

Sky-scraper

Humans

Fingernail

Pin-head

Dust

Bacteria

Virus

Atom

Atomic nucleus

Radio window

Optical window

Opacity
(percent)

Atmosphere
is opaque

Atmosphere
is opaque

100 m 1 m 1 cm
10 m 10 cm 100 μ m 1 μ m

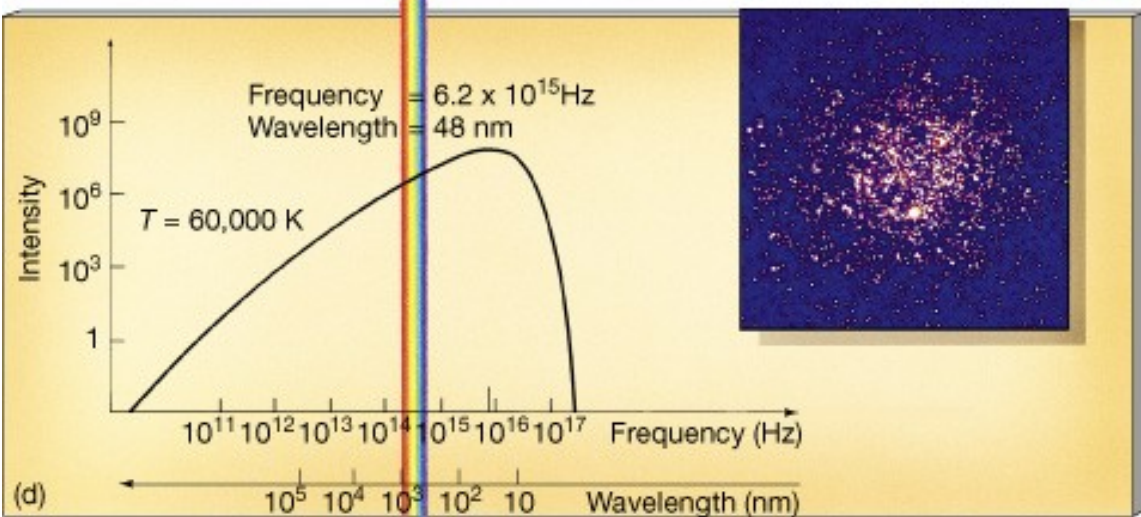
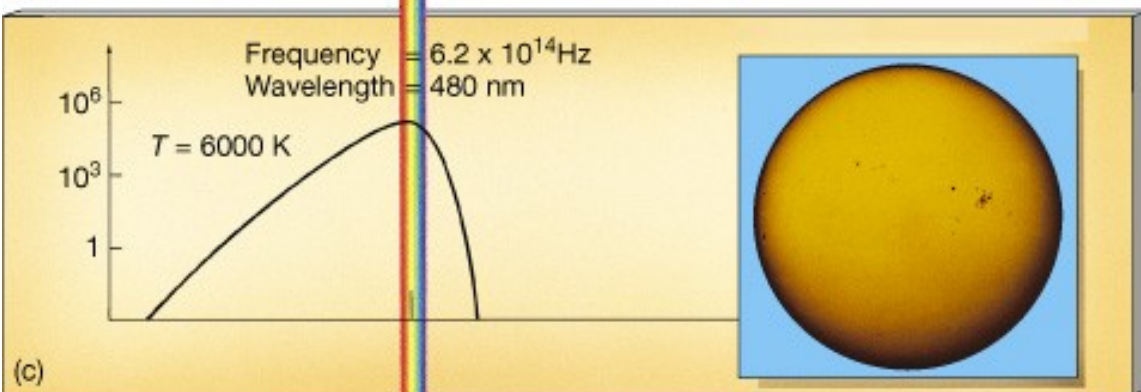
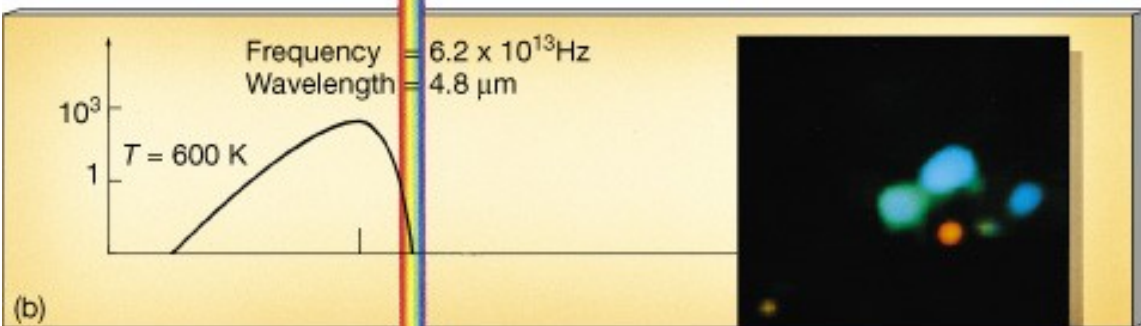
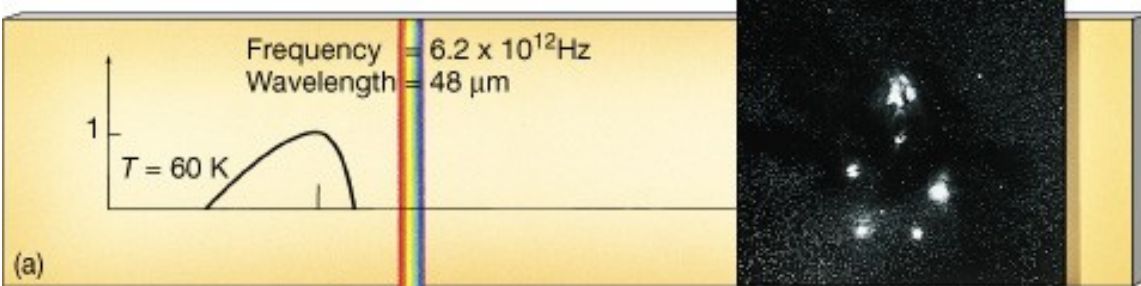
Radiação de Corpo Negro: O Termômetro astronômico

Lei de Wien

Lei de Steffan-Boltzmann

Luminosidade total

Infrared Visible spectrum Ultraviolet



Lei de Wien:

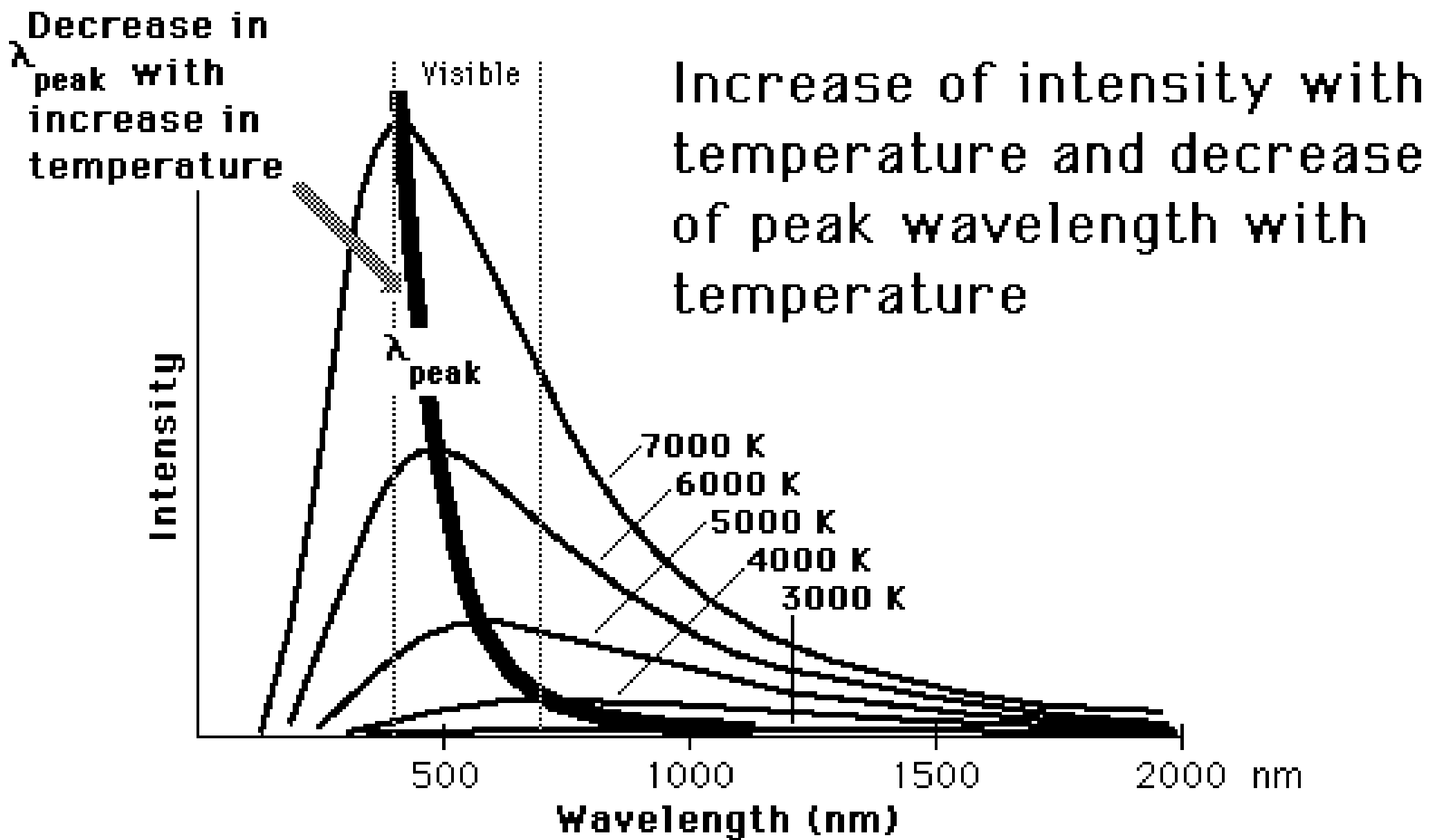
$$\lambda T = 2.9 \times 10^{-3} \text{ A K}$$

Lei de Steffan Boltzmann:

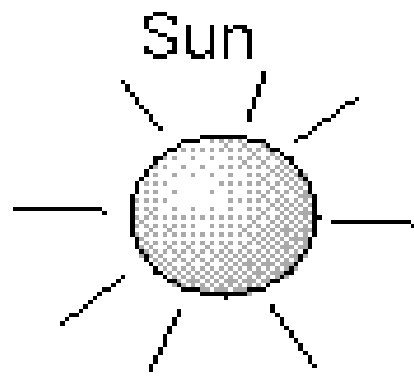
$$F = \sigma T^4$$

$$L = A * F$$

$$A = 4\pi R^2$$



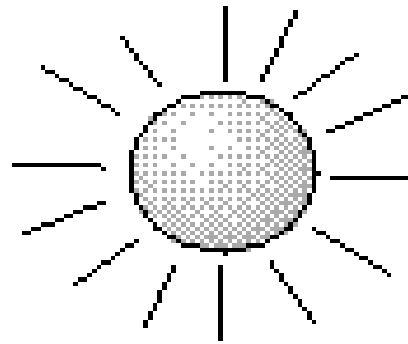
Luminosity is proportional to *fourth* power of temperature.



Sun

6000 K

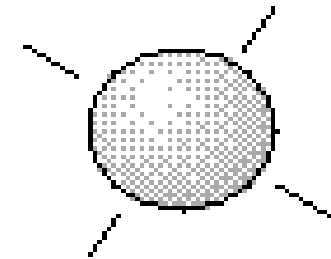
$$L = 1$$



12,000 K

$$L = 16$$

$$\left(\frac{12,000}{6000}\right)^4 = 2^4$$

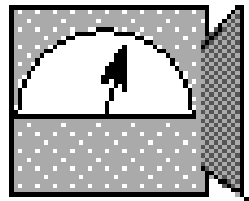


2000 K

$$L = \frac{1}{81}$$

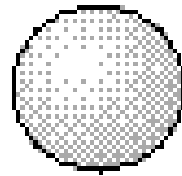
$$\left(\frac{2000}{6000}\right)^4 = (1/3)^4$$

Luminosity is also proportional to the surface area.



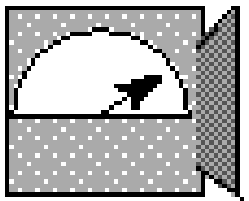
100 photons

Sun

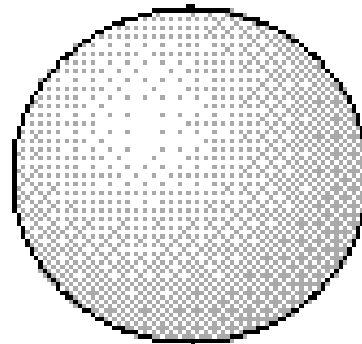


$R = 1 \text{ sun}$

$$\text{surface area} = 4\pi R^2$$



400 photons



$R = 2 \text{ sun}$

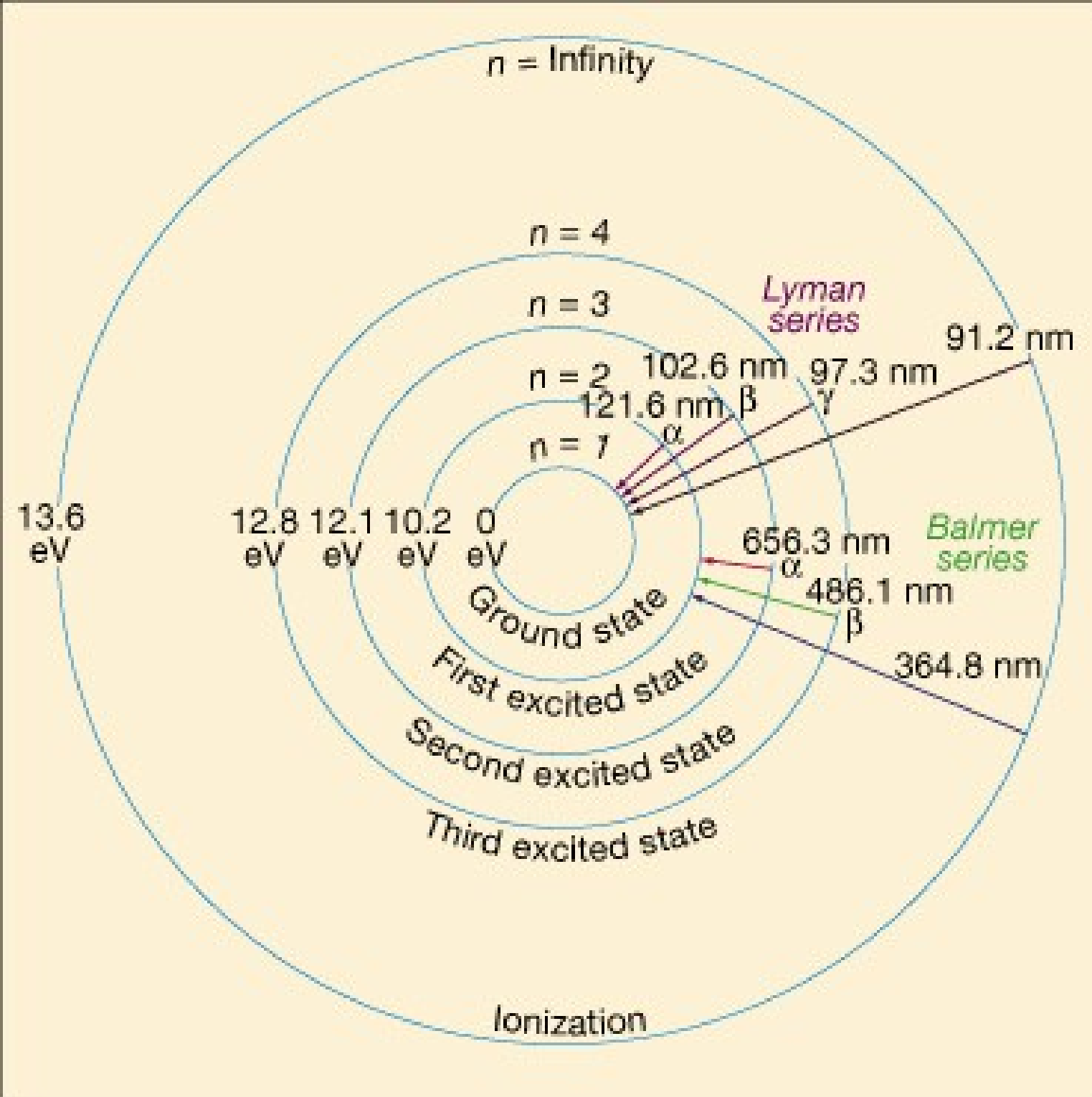
$$\frac{\text{star surface area}}{\text{Sun surface area}} = \left(\frac{2}{1}\right)^2$$

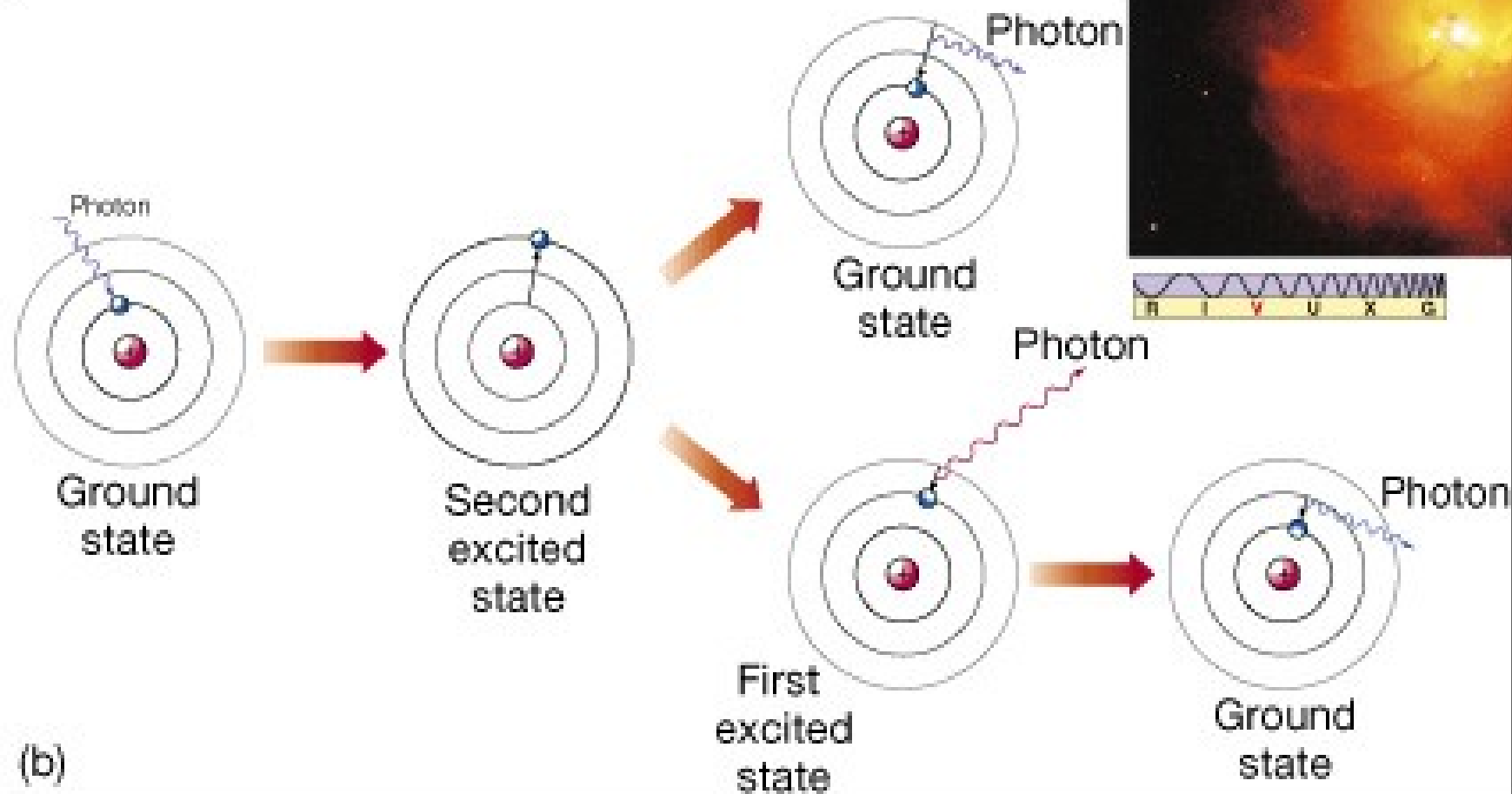
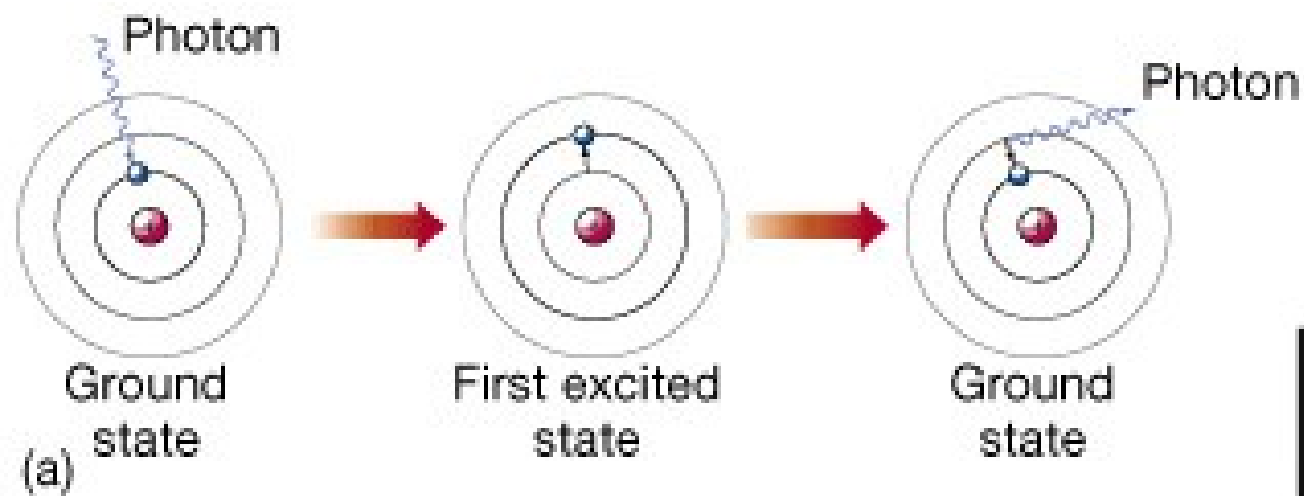
ESPECTROSCOPIA:

Emissão e absorção de fótons

O átomo do hidrogênio

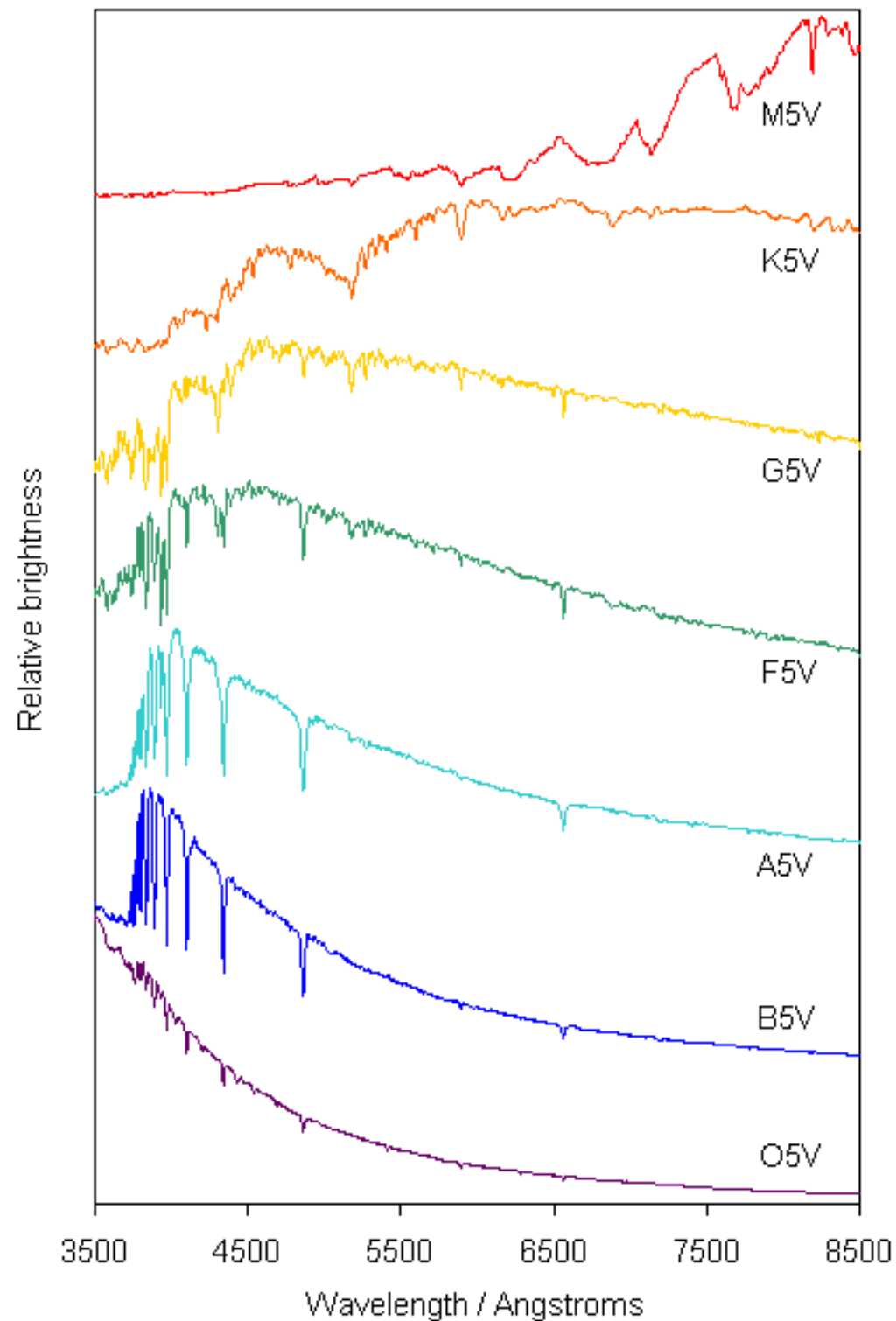
O modelo de Bohr

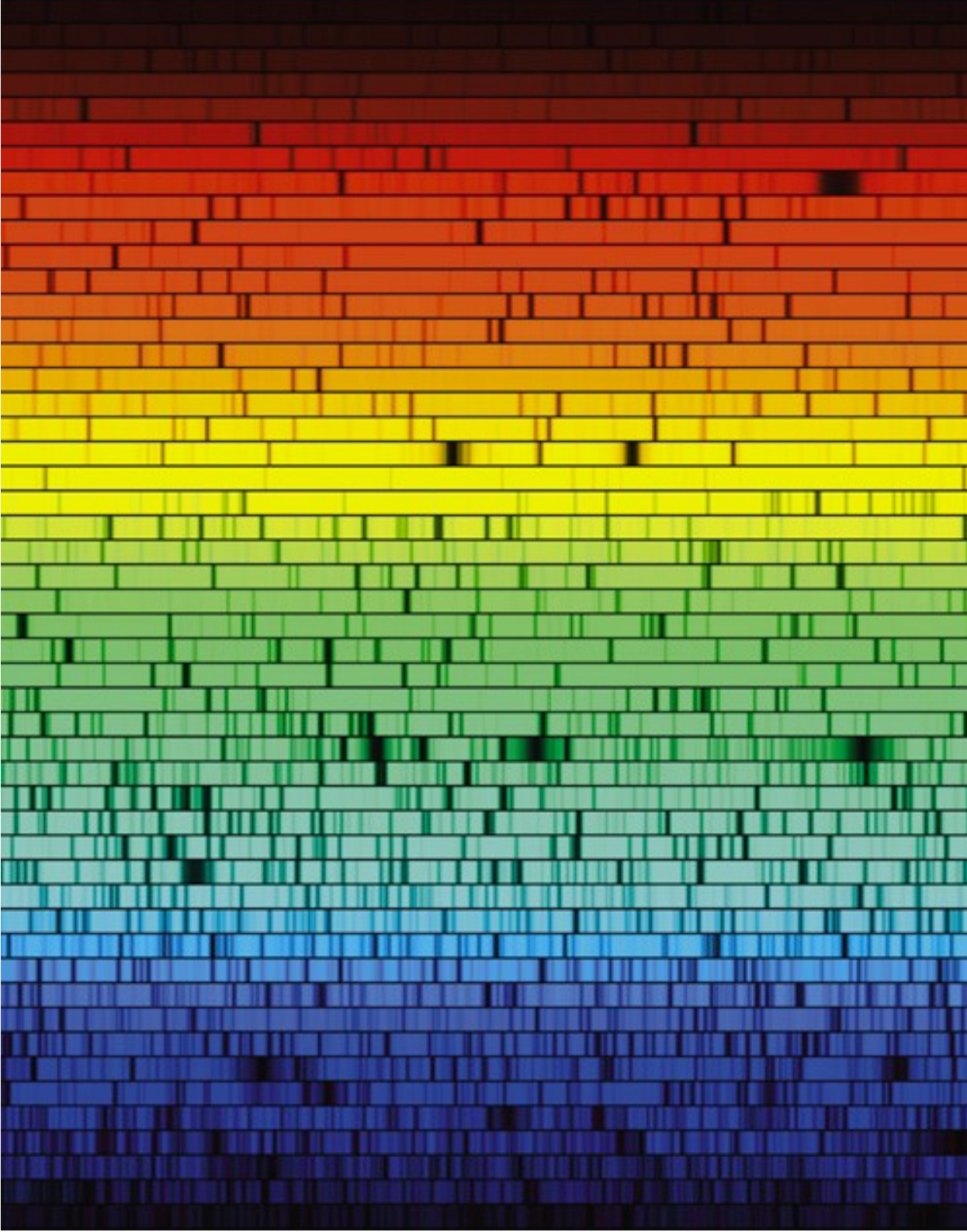




Espectro de algumas estrelas “comuns”

Notem para as estrelas de O5
até G5 a presença de linhas de
hidrogênio.

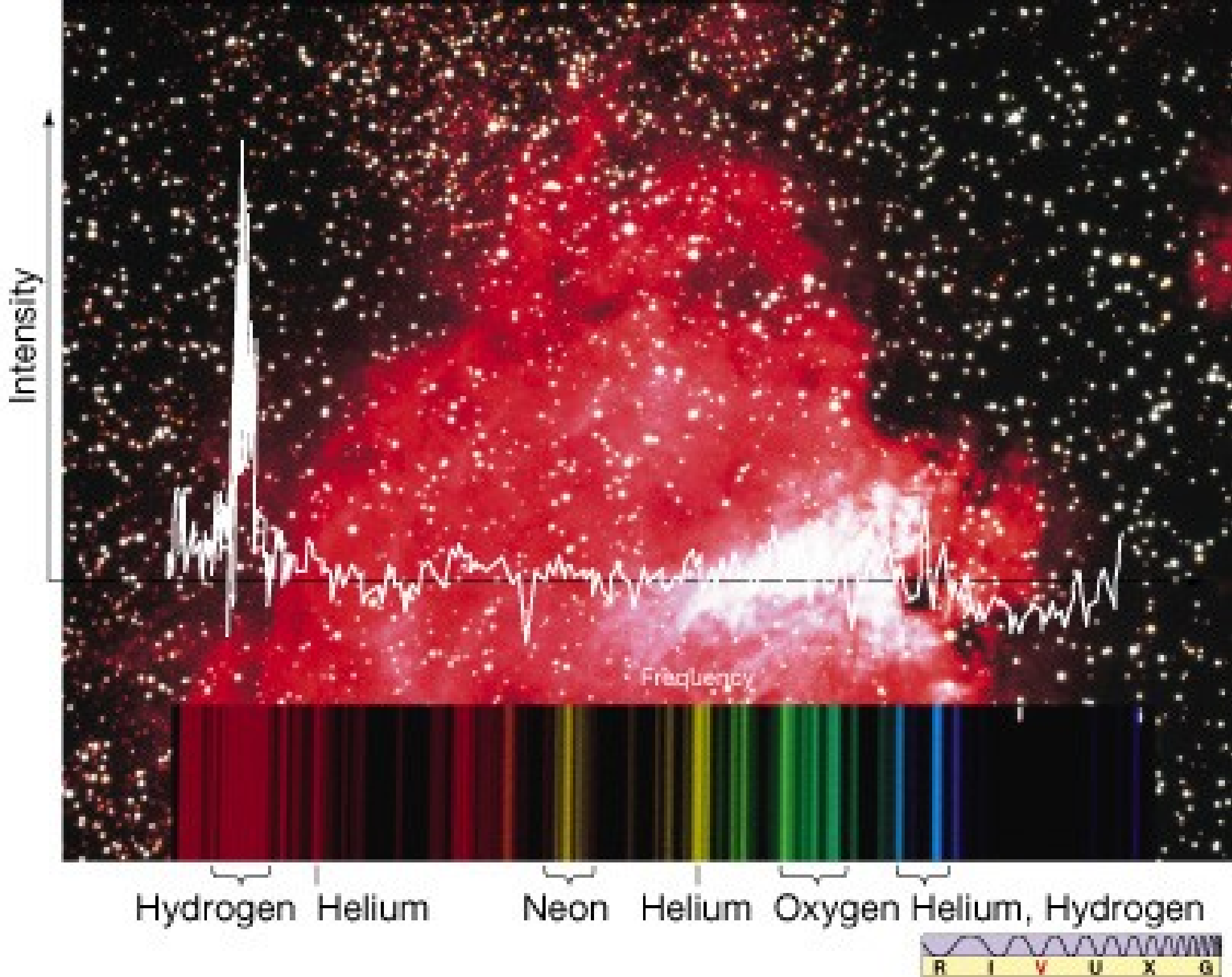




Espectro do Sol

Esta é uma imagem do espectro solar. As cores indicam a região do espectro, e apesar de falsas indicam corretamente a posição no espectro visível.

As manchas escuras são as linhas de absorção (leis de Kirchhoff logo adiante)

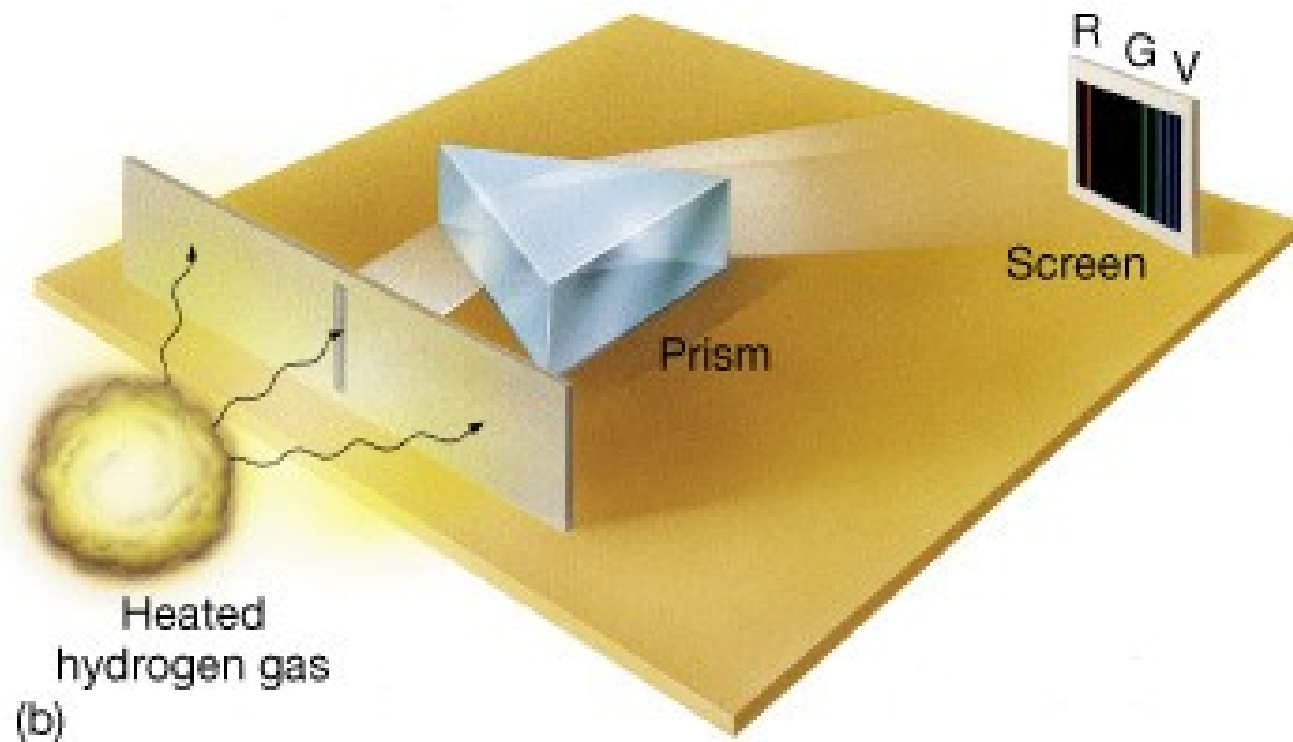
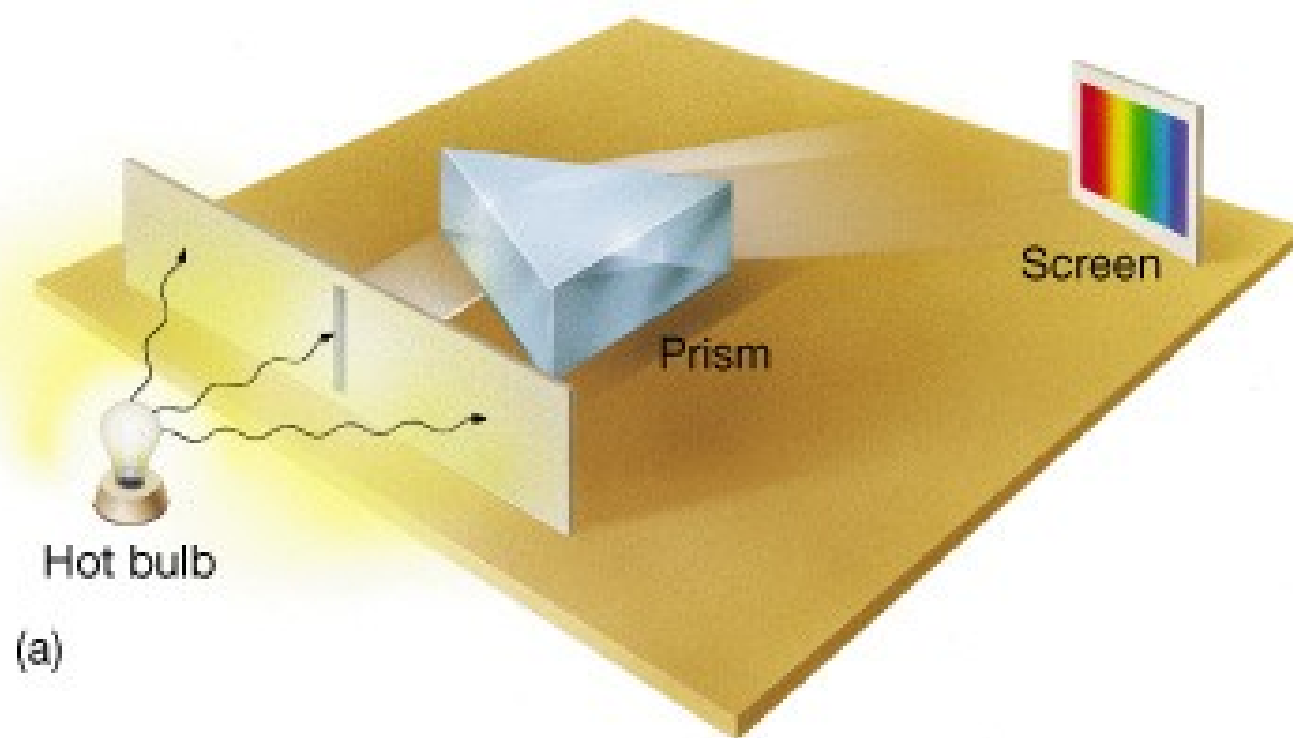


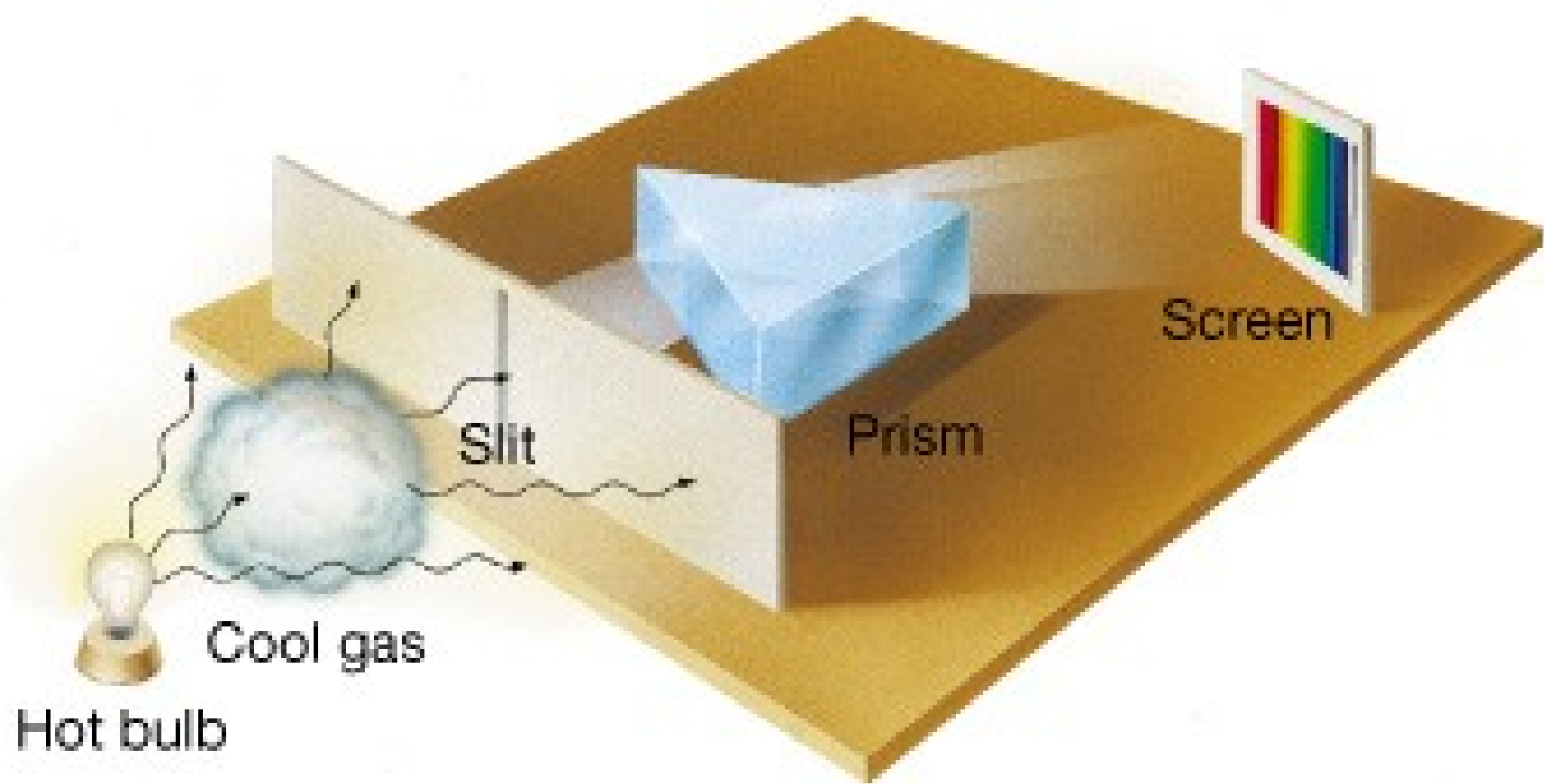
Leis de Kirchhoff:

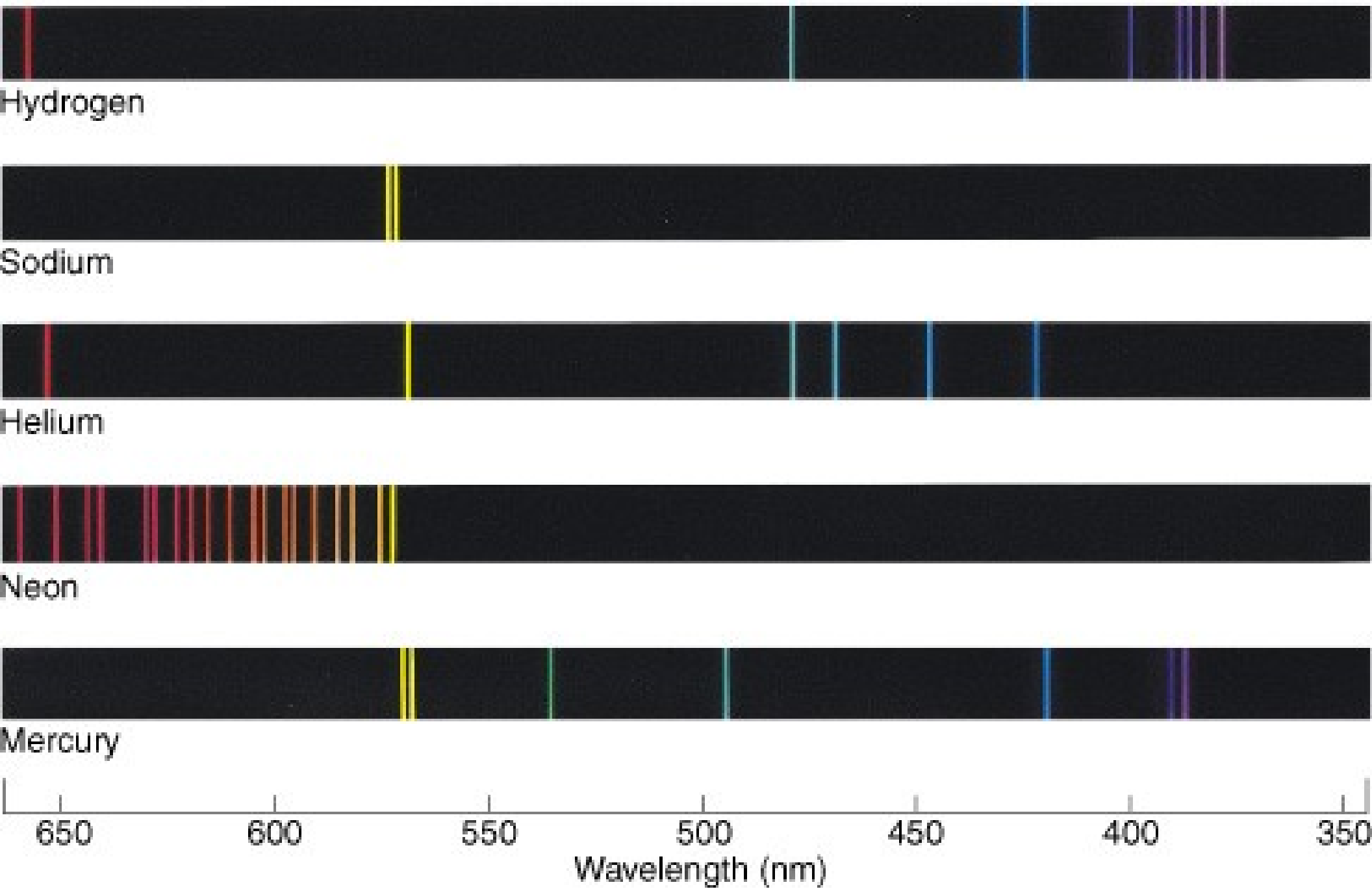
1 – contínuo – corpo denso e opaco

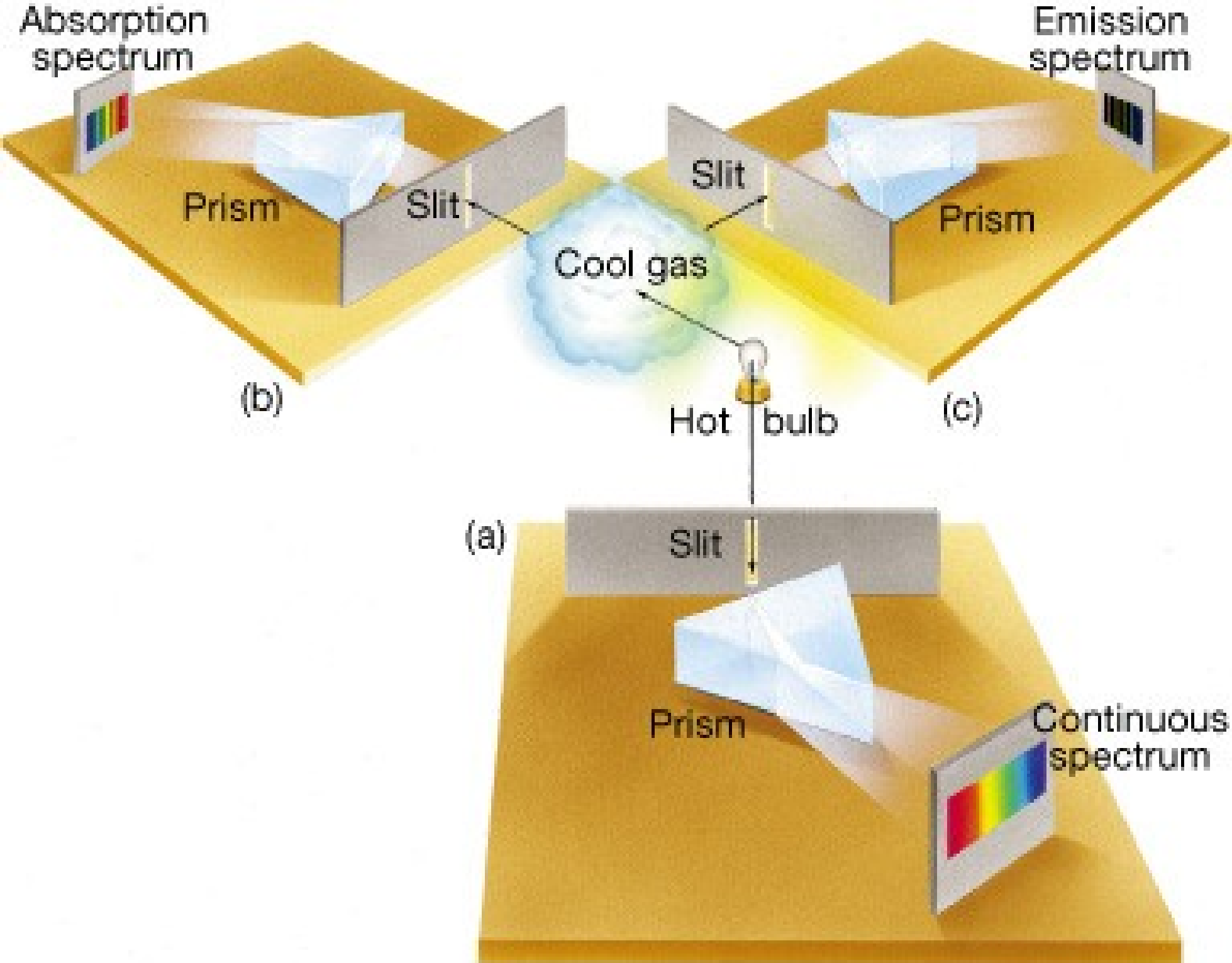
2 – emissão – gás fino

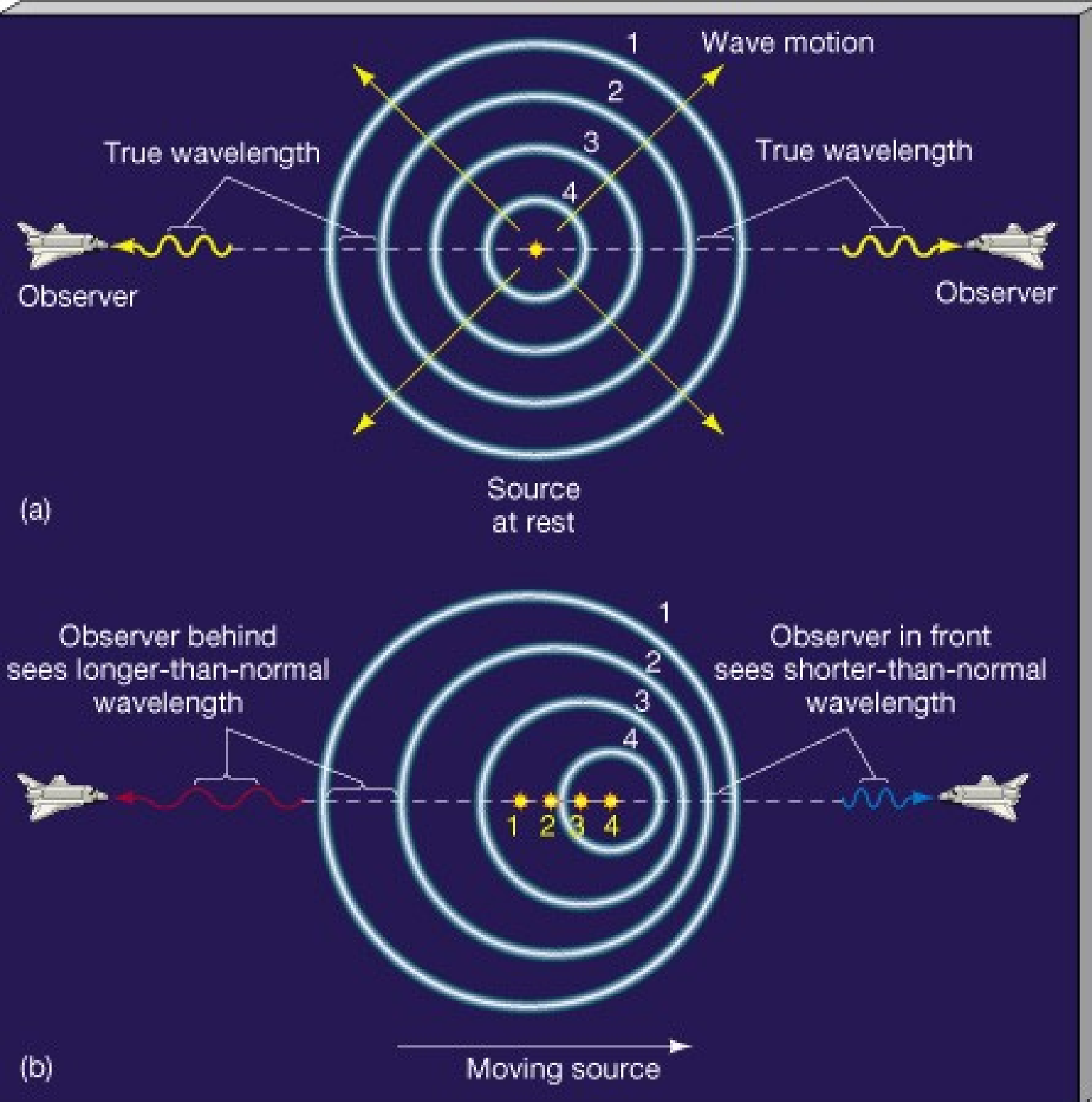
3 – absorção - gás fino iluminado por trás por algum outro objeto











For waves that travel through a medium (sound, ultrasound, etc...) the relationship between observed frequency f' and emitted frequency f is given by:

$$f' = \left(\frac{v}{v \pm v_s} \right) f$$

where

v , is the speed of waves in the medium (in air at T degrees Celsius, this is $332 + 0.59T$ m/s)

v_s , is the velocity of the source (the thing emitting the sound)

For waves that travel at the speed of light, such as radio waves, the relationship between observed frequency f' and emitted frequency f is given by:

Change in frequency Observed frequency

$$\Delta f = \frac{fv}{c} = \frac{v}{\lambda}$$

$$f' = f + \frac{fv}{c}$$

where

f , is the transmitted frequency

v , is the velocity of the transmitter relative to the receiver in meters/second: positive when moving towards one another, negative when moving away

c , is the speed of light in a vacuum (3×10^8) m/s

λ , is the wavelength of the transmitted wave subject to change

For motion along the line of sight

If the observer and the source are moving directly away from each other with velocity v , the observed frequency f_o , is different from the frequency of the source f_e , as

$$f_o = \sqrt{\frac{1-v/c}{1+v/c}} f_e$$

where c , is the speed of light.

The corresponding wavelengths are related by

$$\lambda_o = \sqrt{\frac{1+v/c}{1-v/c}} \lambda_e$$

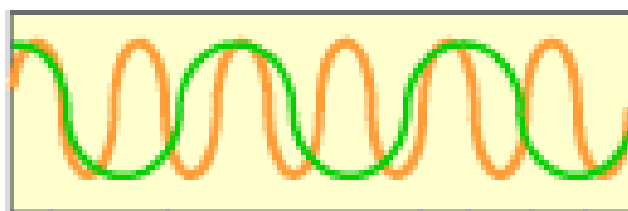
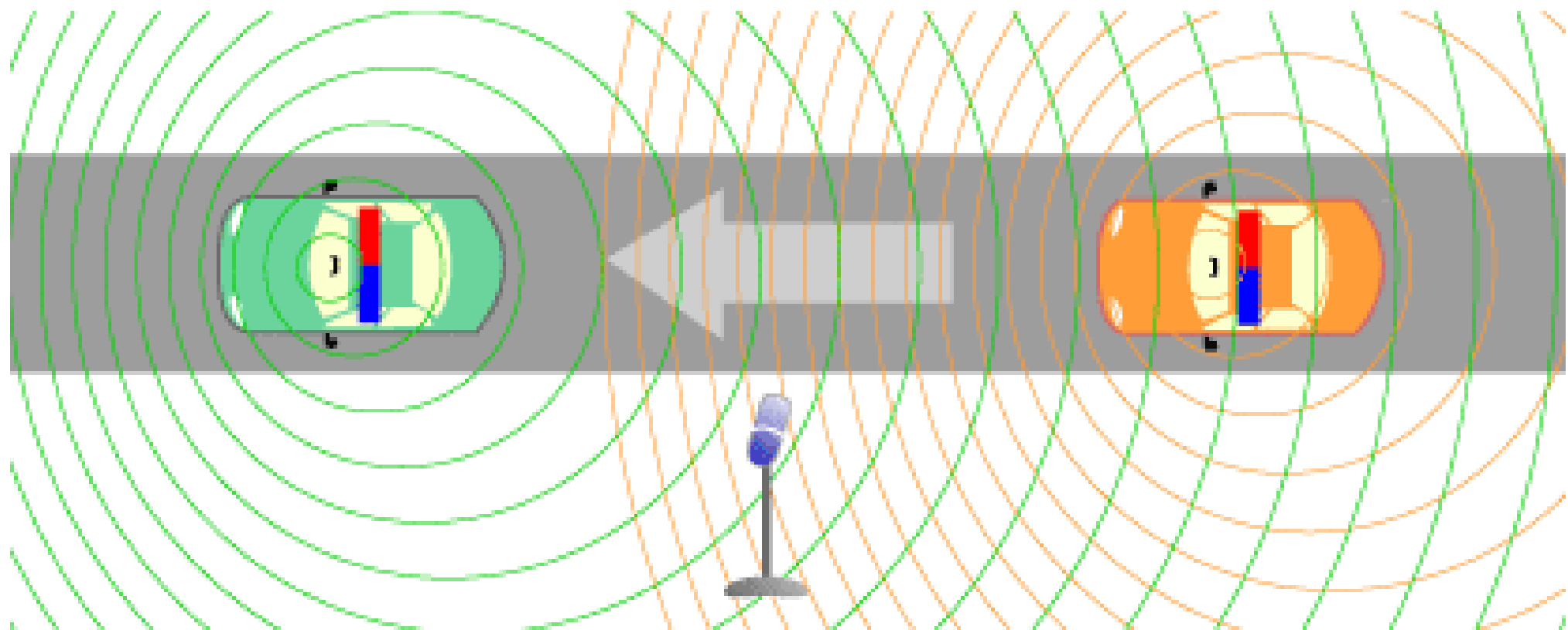
and the resulting redshift z , can be written as

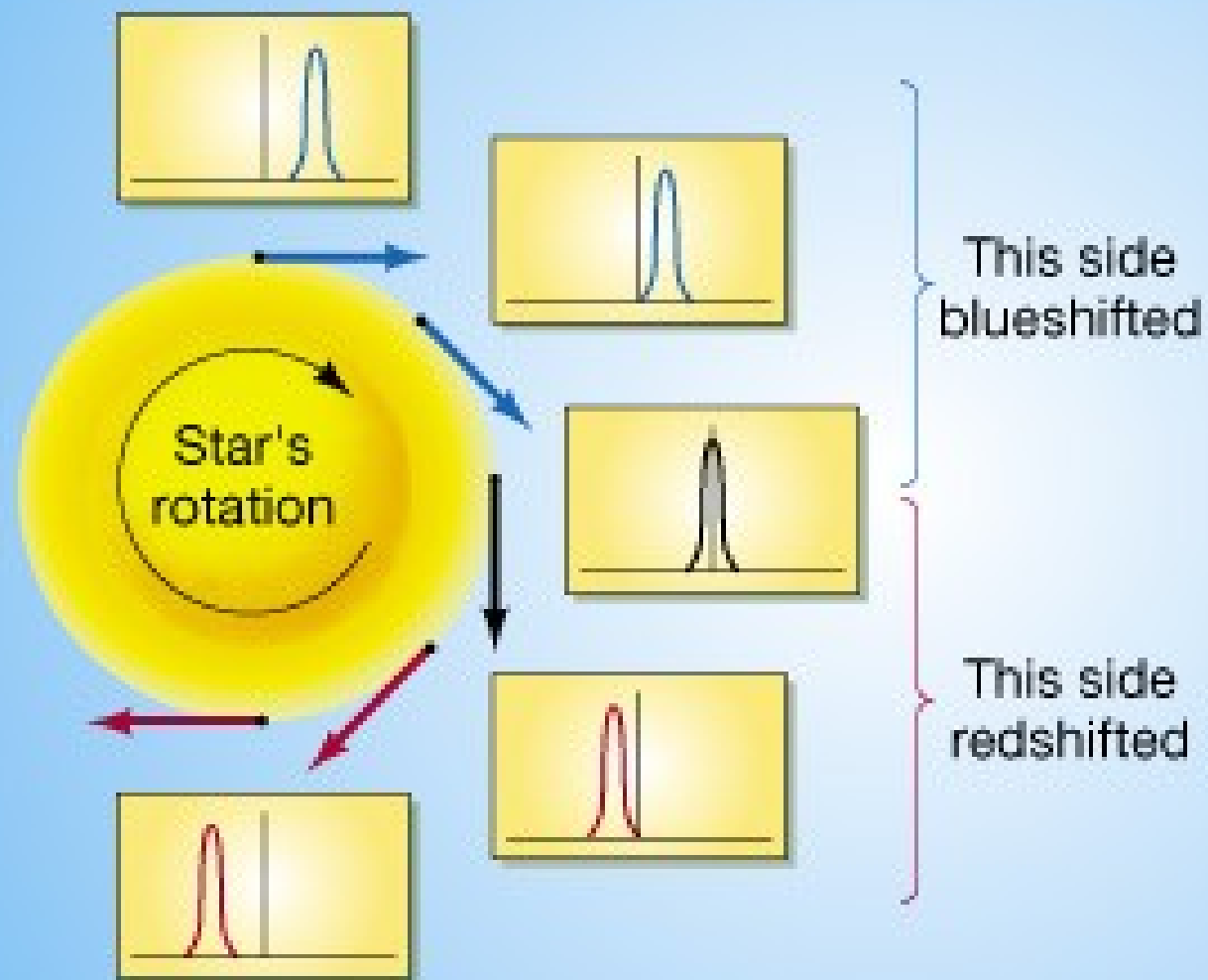
$$z + 1 = \frac{\lambda_o}{\lambda_e} = \sqrt{\frac{1+v/c}{1-v/c}}$$

In the non-relativistic limit, i.e. when $v \ll c$, the approximate expressions are:

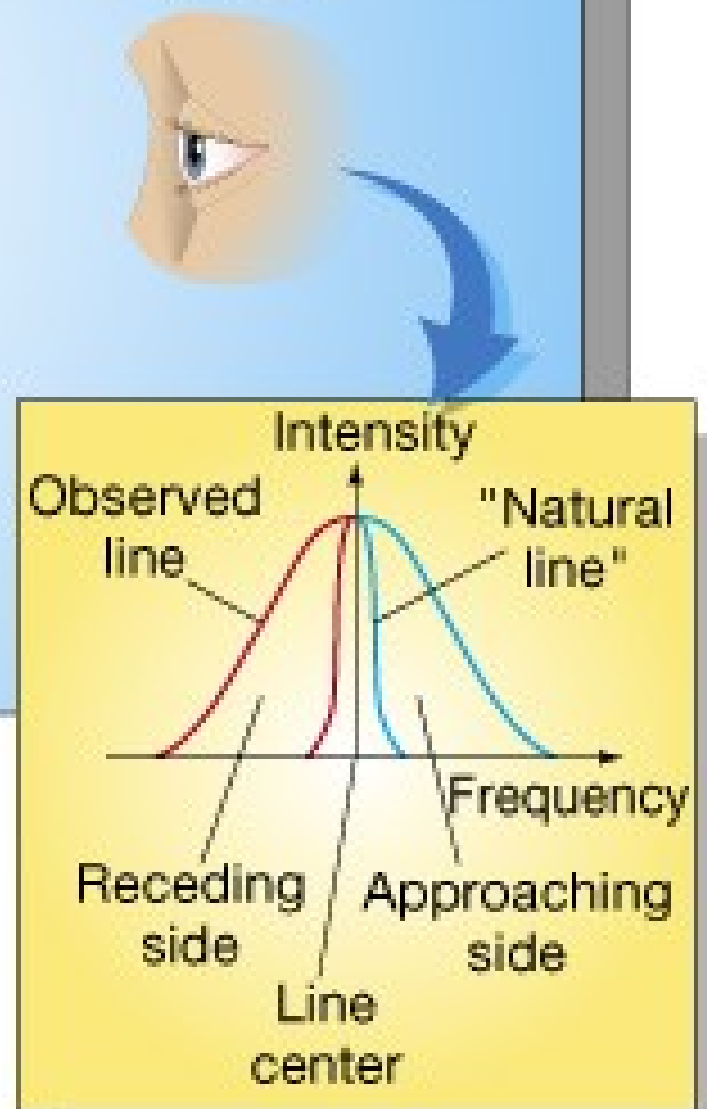
$$\frac{\Delta f}{f} \simeq -\frac{v}{c} \quad \frac{\Delta \lambda}{\lambda} \simeq \frac{v}{c} \quad z \simeq \frac{v}{c}$$

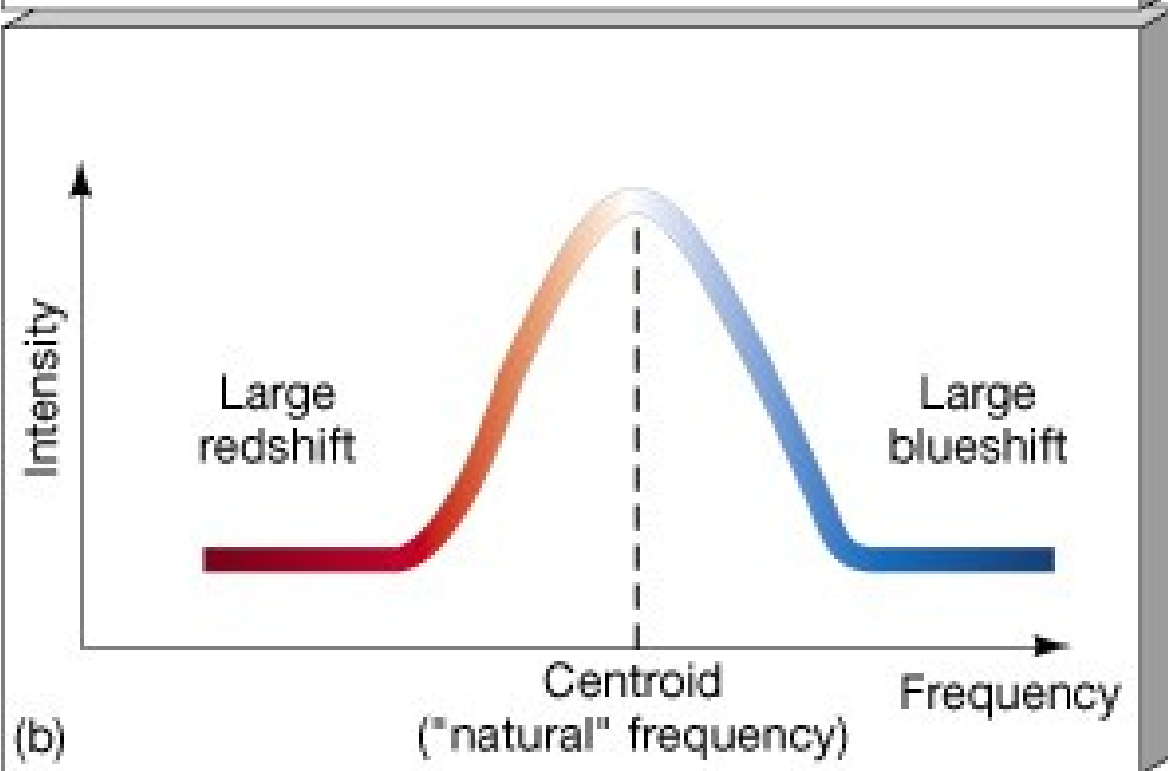
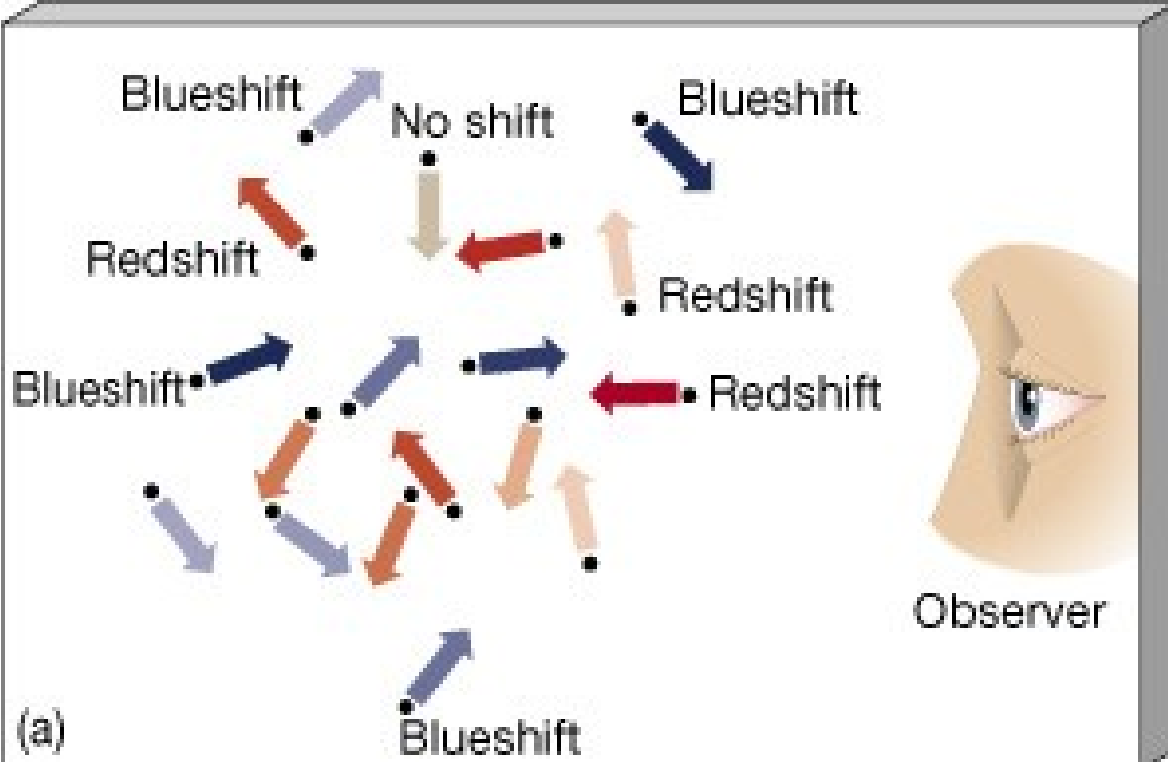
Note: In all the expressions in this section it is assumed that the observer and the source are moving away from each other. If they are moving towards each other, v , should be taken negative.





Observer
sees:

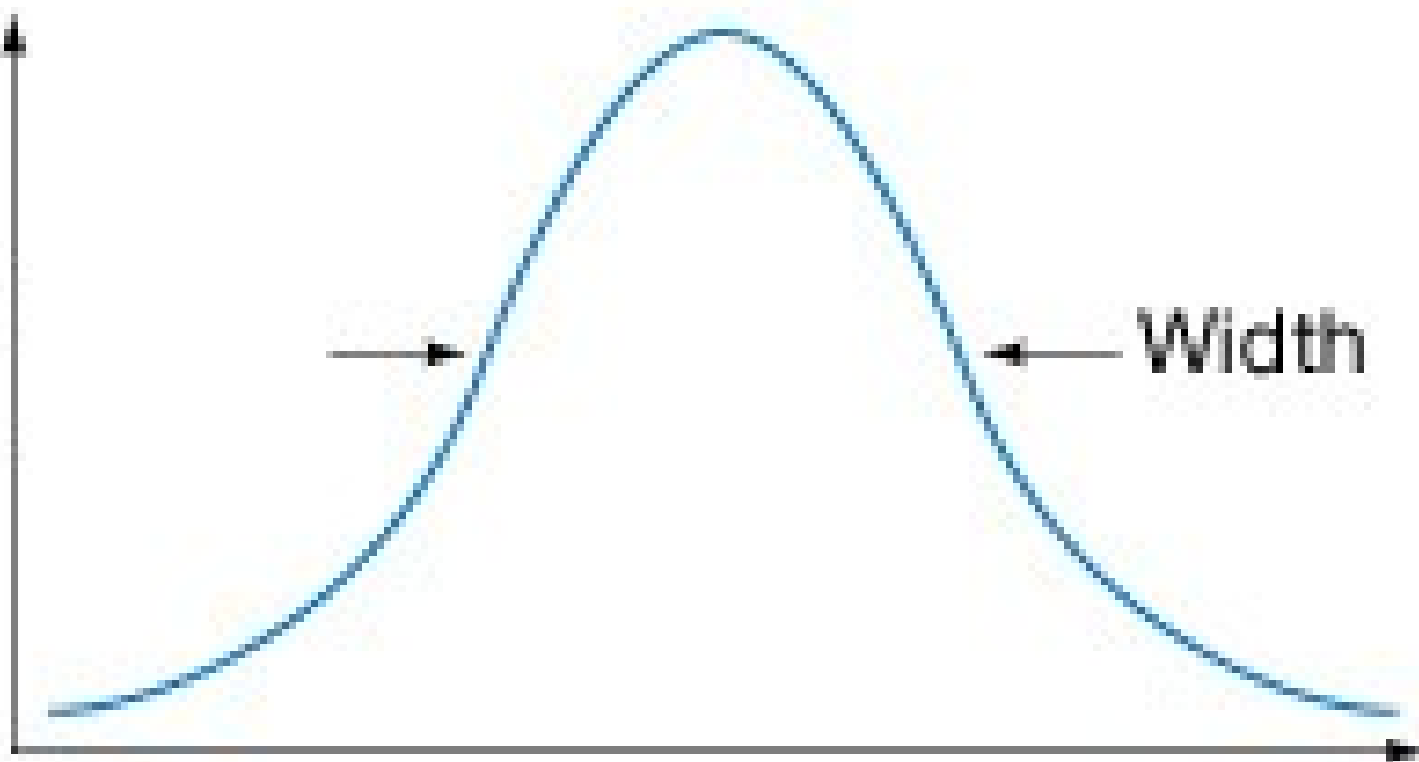




(a)



Intensity



Width

Frequency

(b)